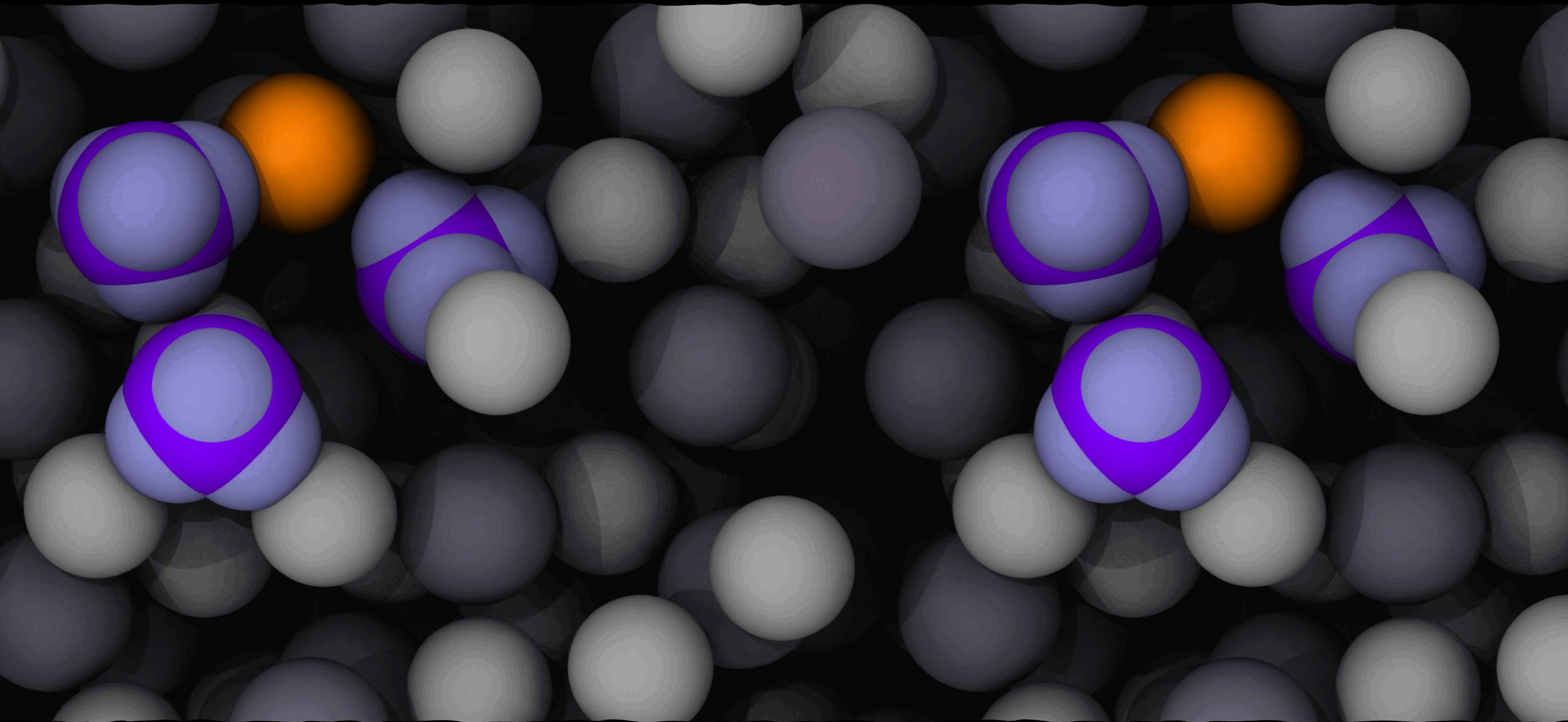


Predicting Properties from First-Principles II: *Total Energy and Force Applications*



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Nistha Sheth*

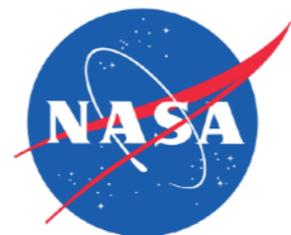
Funding:



ASTROBIOLOGY at NASA
LIFE IN THE UNIVERSE



U.S. DEPARTMENT OF
ENERGY

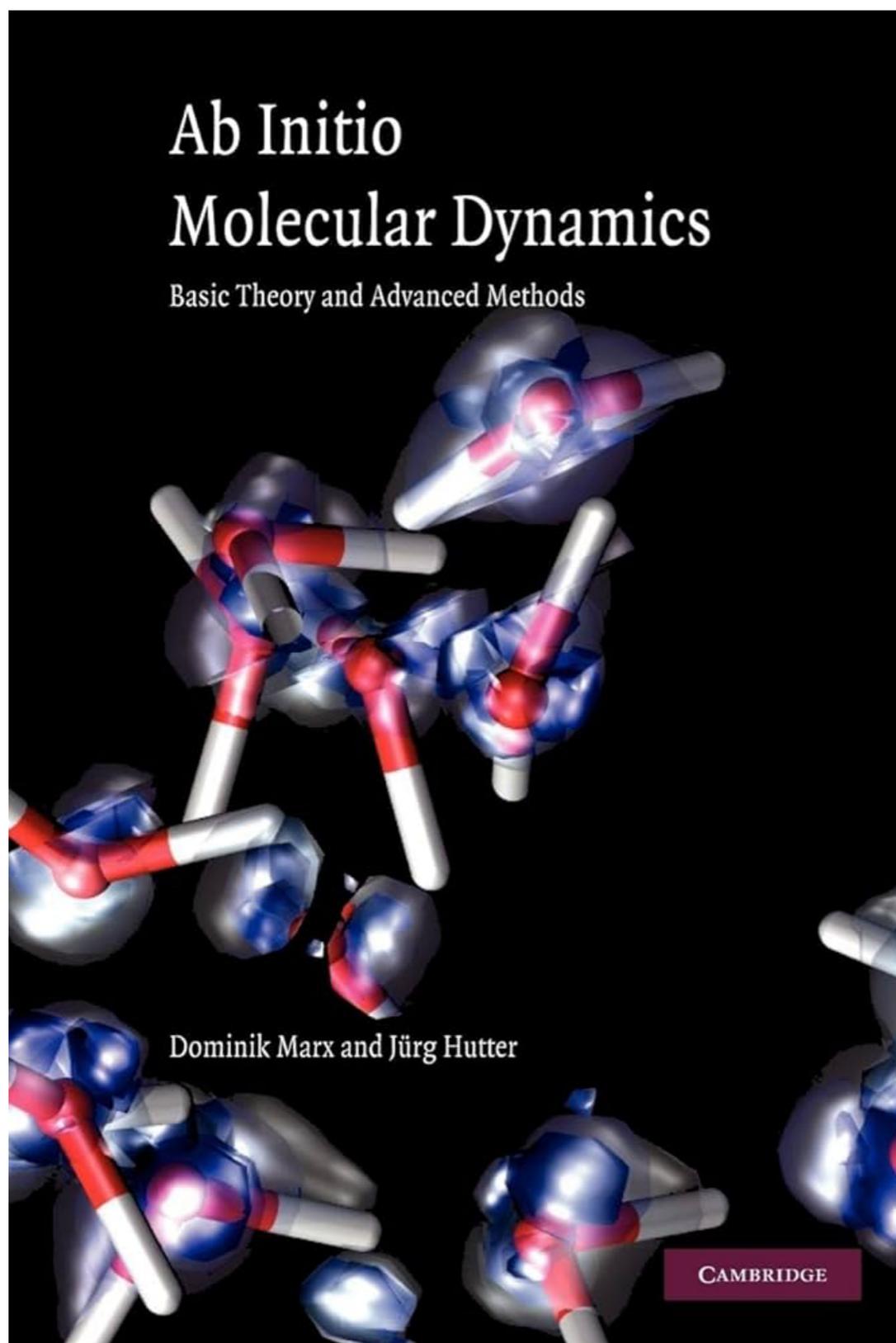


Planetary Science
Early Career Award

XSEDE

Extreme Science and Engineering
Discovery Environment

Good References to Learn More



The Journal of Chemical Physics

ARTICLE

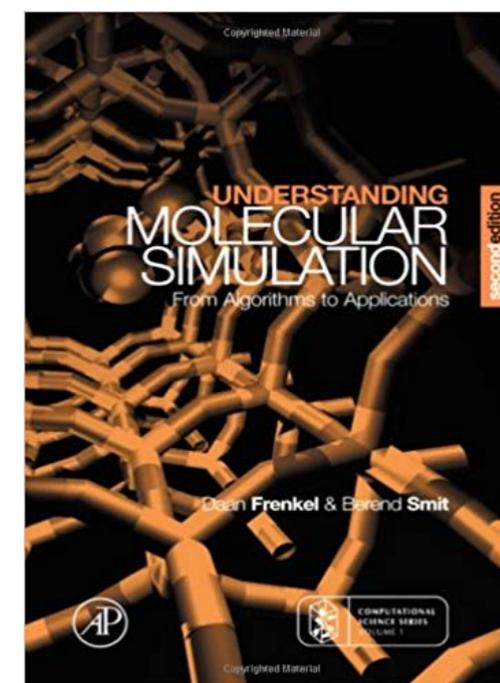
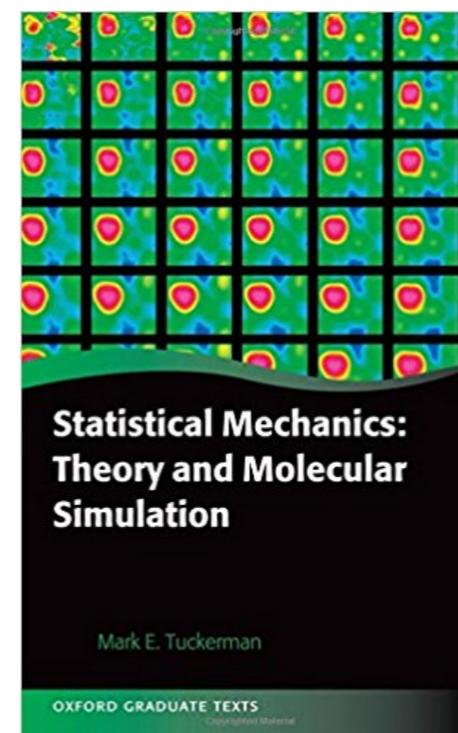
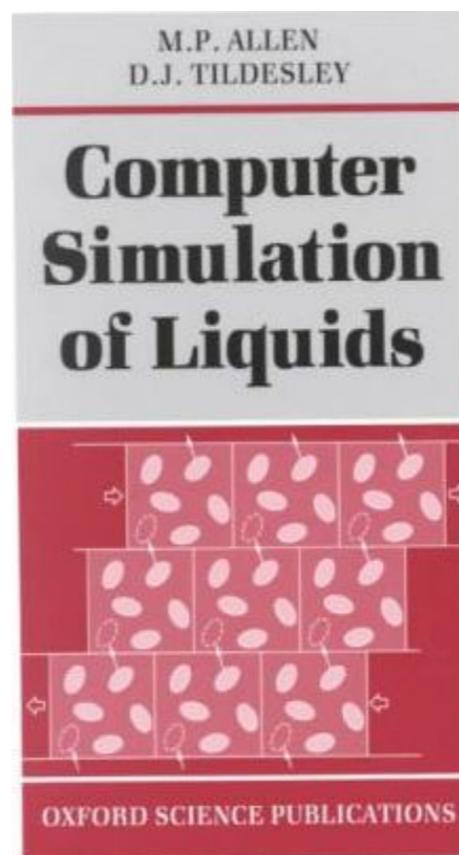
scitation.org/journal/jcp

CP2K: An electronic structure and molecular dynamics software package - Quickstep: Efficient and accurate electronic structure calculations

Cite as: J. Chem. Phys. 152, 194103 (2020); doi: [10.1063/5.0007045](https://doi.org/10.1063/5.0007045)
Submitted: 10 March 2020 • Accepted: 22 April 2020 •
Published Online: 19 May 2020

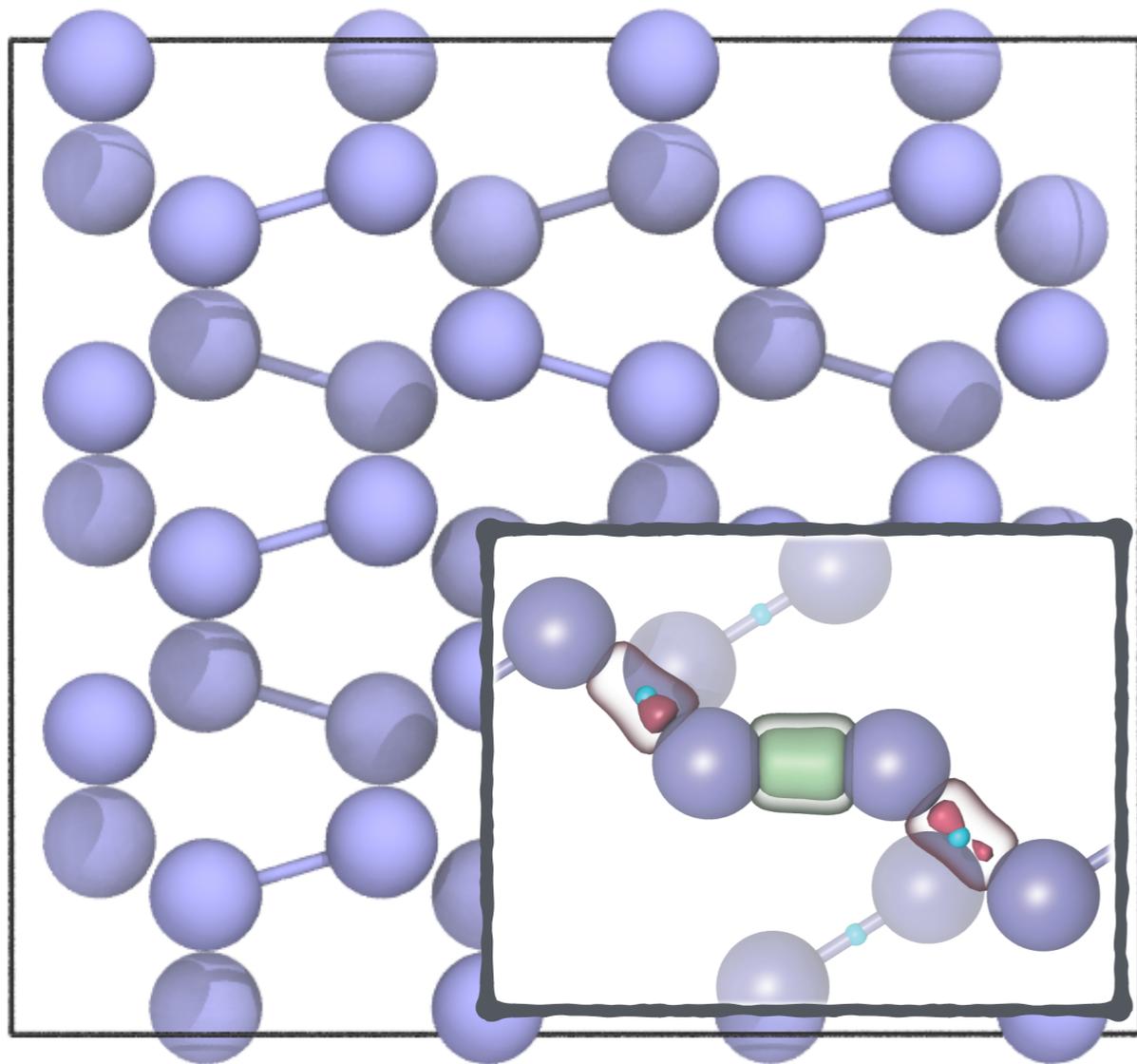
Thomas D. Kühne,^{1,a)}  Marcella Iannuzzi,²  Mauro Del Ben,³ Vladimir V. Rybkin,²  Patrick Seewald,² Frederick Stein,² Teodoro Laino,⁴  Rustam Z. Khaliullin,⁵  Ole Schütt,⁶ Florian Schiffmann,⁷  Dorothea Golze,⁹  Jan Wilhelm,⁹ Sergey Chulkov,¹⁰ Mohammad Hossein Bani-Hashemian,¹¹  Valéry Weber,⁴ Urban Borštnik,¹² Mathieu Taillefumier,¹³ Alice Shoshana Jakobovits,¹³  Alfio Lazzaro,¹⁴ Hans Pabst,¹⁵ Tiziano Müller,² Robert Schade,¹⁶  Manuel Guidon,² Samuel Andermatt,¹¹ Nico Holmberg,¹⁷ Gregory K. Schenter,¹⁸  Anna Hehn,² Augustin Bussy,² Fabian Belleflamme,² Gloria Tabacchi,¹⁹  Andreas Glöb,²⁰ Michael Lass,¹⁶  Iain Bethune,²¹  Christopher J. Mundy,¹⁸  Christian Pleschl,¹⁶  Matt Watkins,¹⁰  Joost VandeVondele,¹³  Matthias Krack,^{22,b)}  and Jürg Hutter^{2,c)}



Why think about total energy?

- Many important materials properties depend on total energy

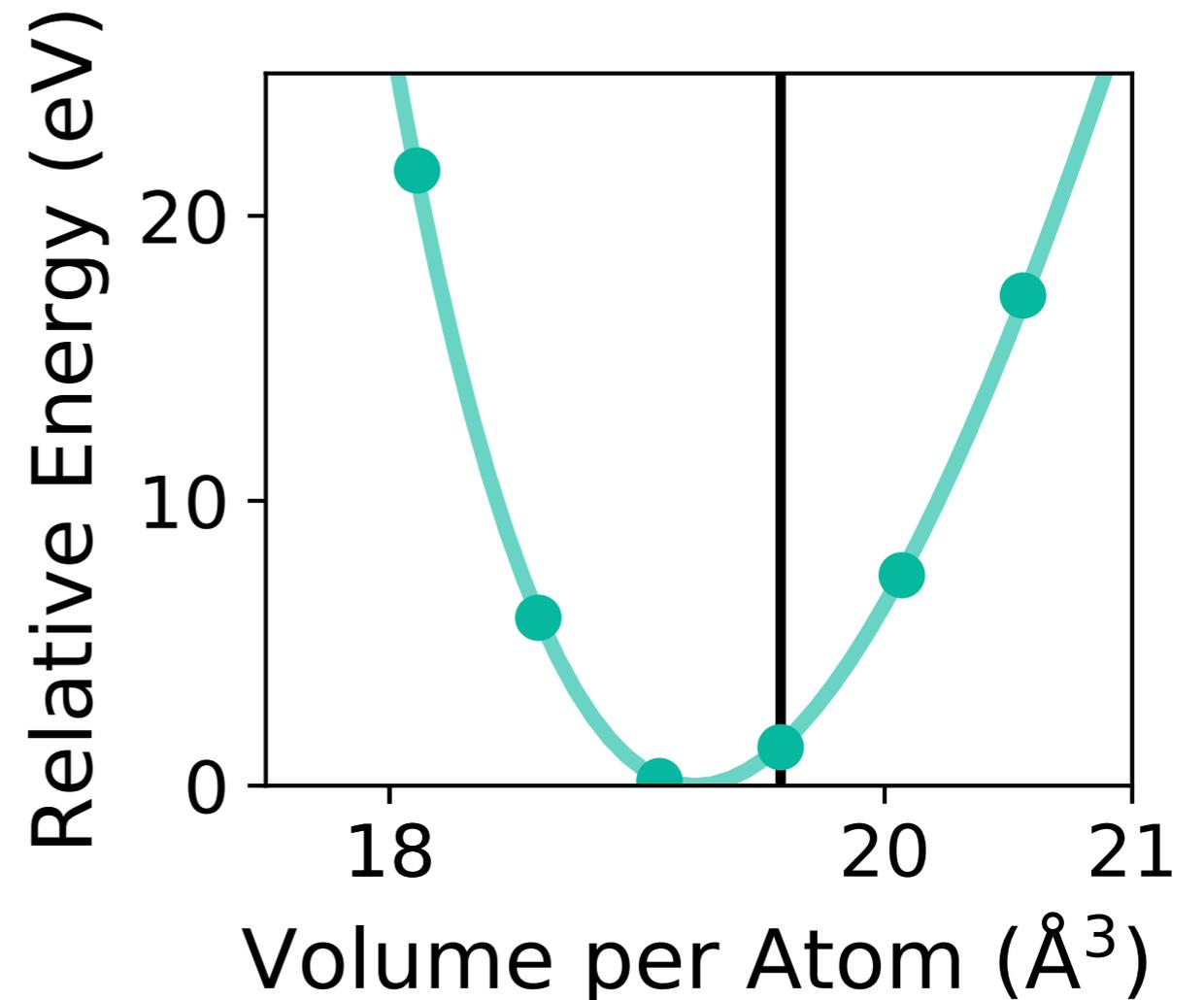
α -Gallium



Murnaghan equation of state:

K_0 = bulk modulus

- Equation of State & Bulk Moduli

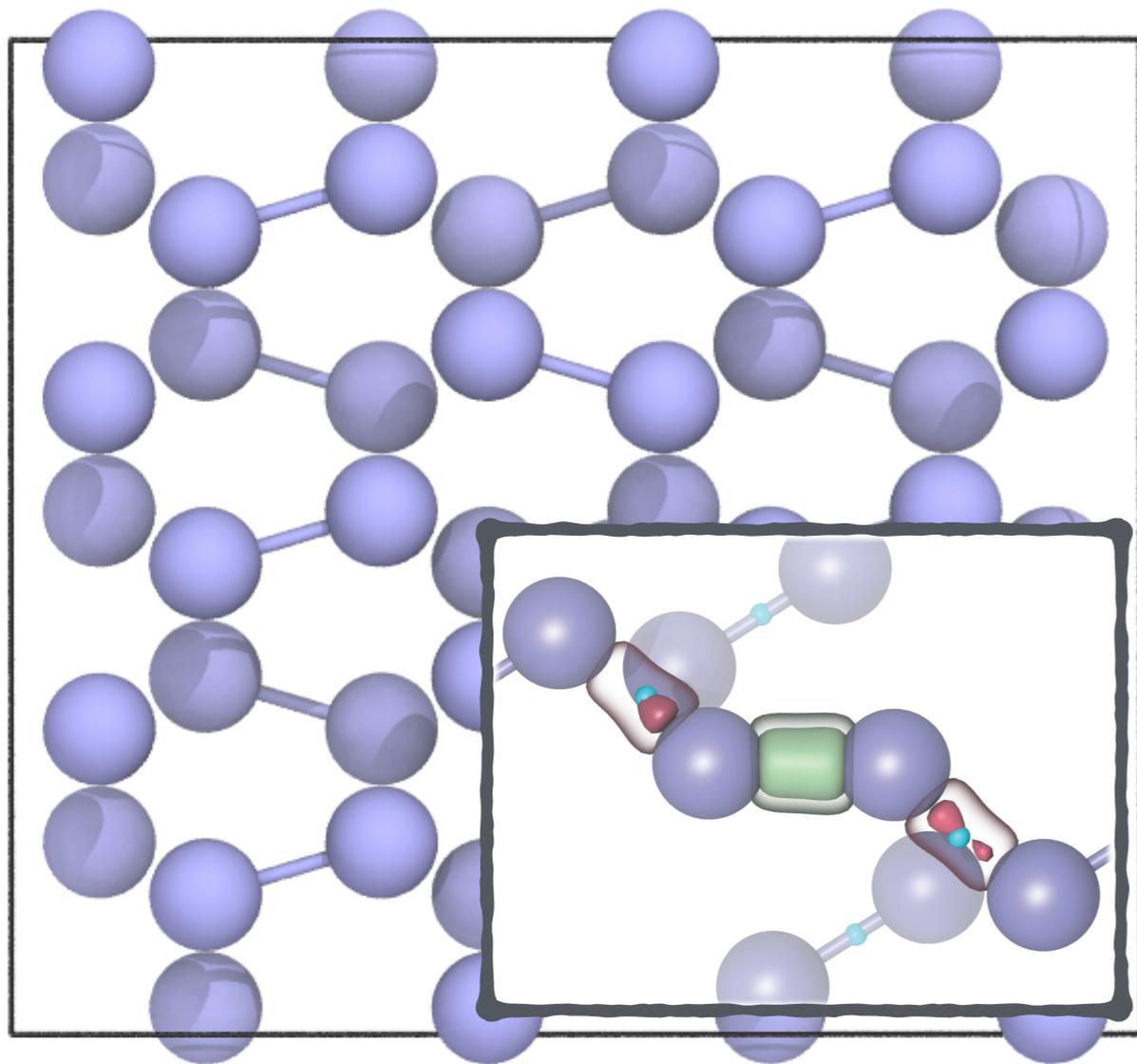


$$E(V) - E_0 = K_0 V_0 \left[\frac{1}{K'_0(K'_0 - 1)} \left(\frac{V}{V_0} \right)^{1-K'_0} + \frac{1}{K'_0} \frac{V}{V_0} - \frac{1}{K'_0 - 1} \right]$$

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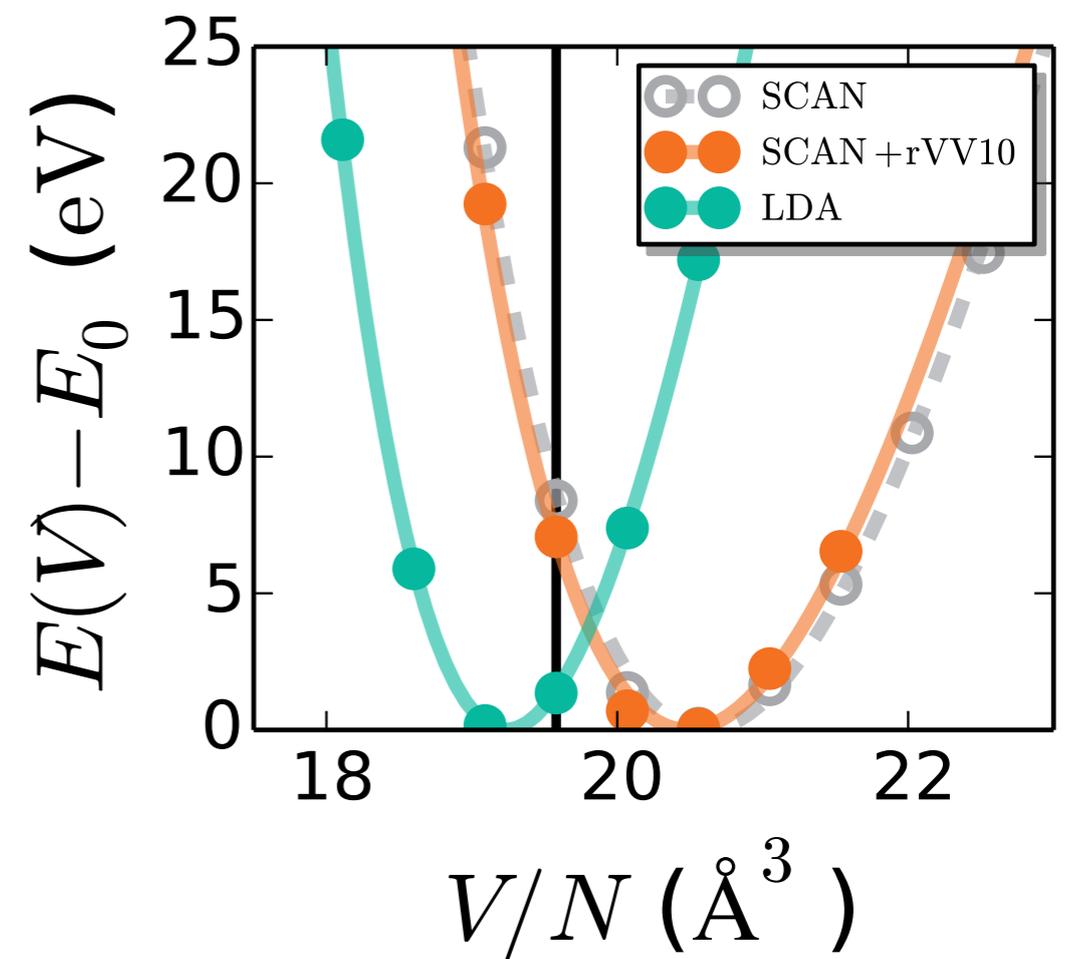


Murnaghan equation of state:

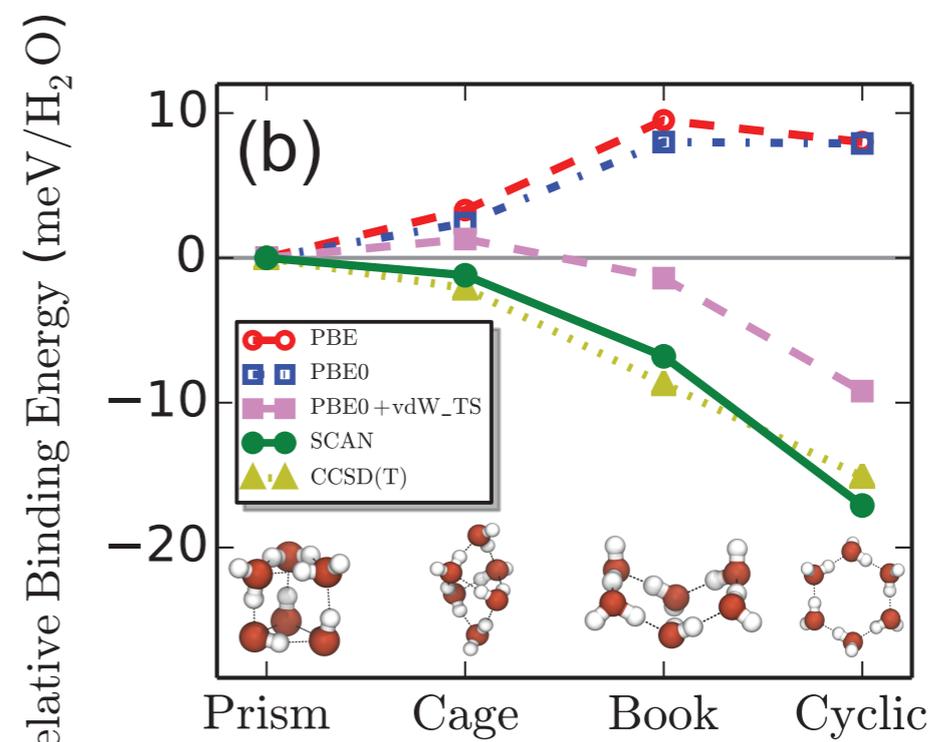
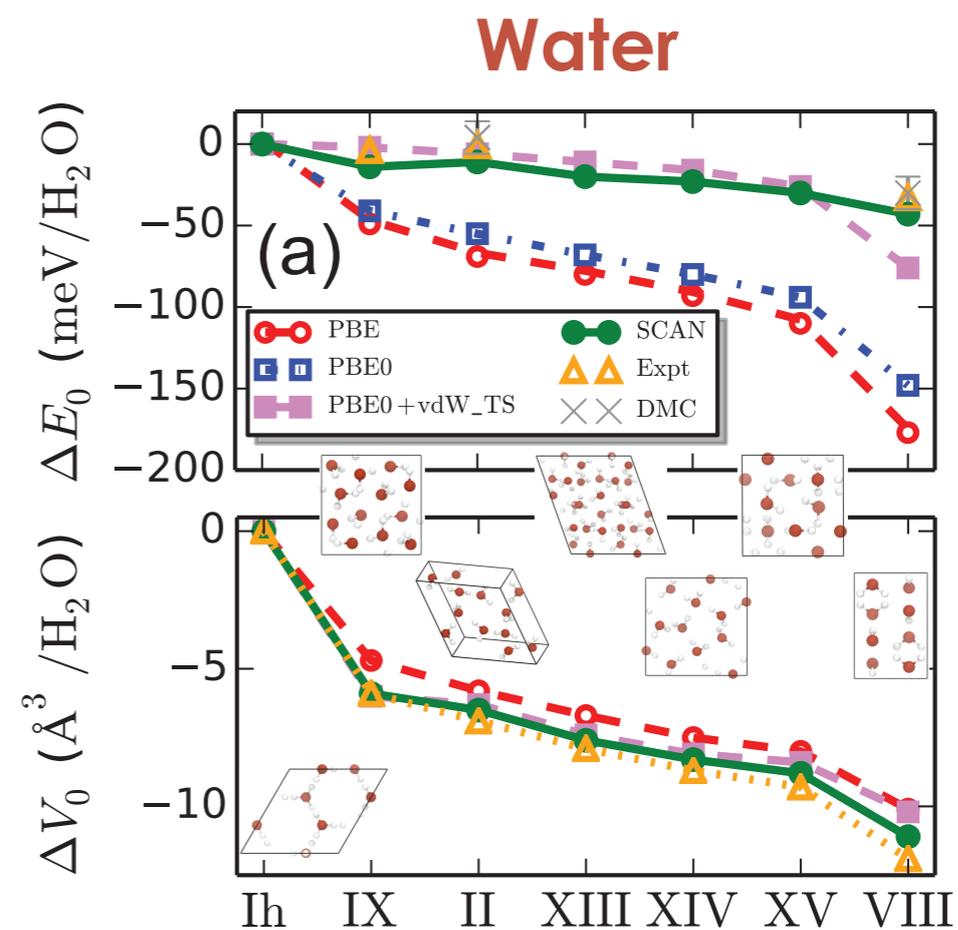
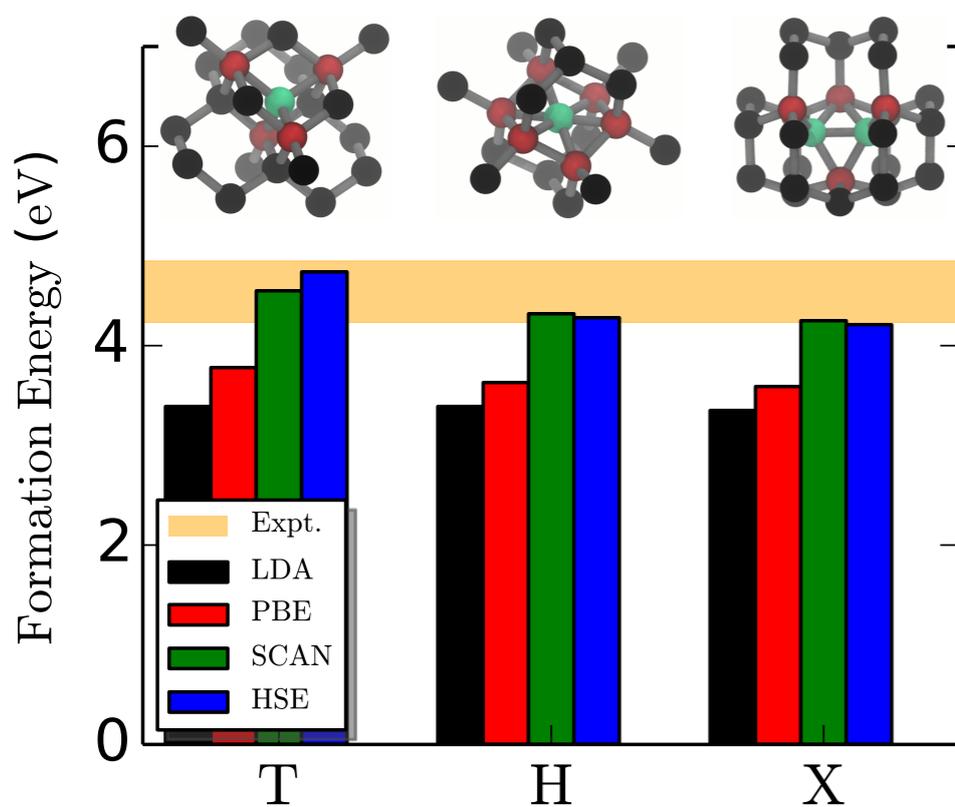
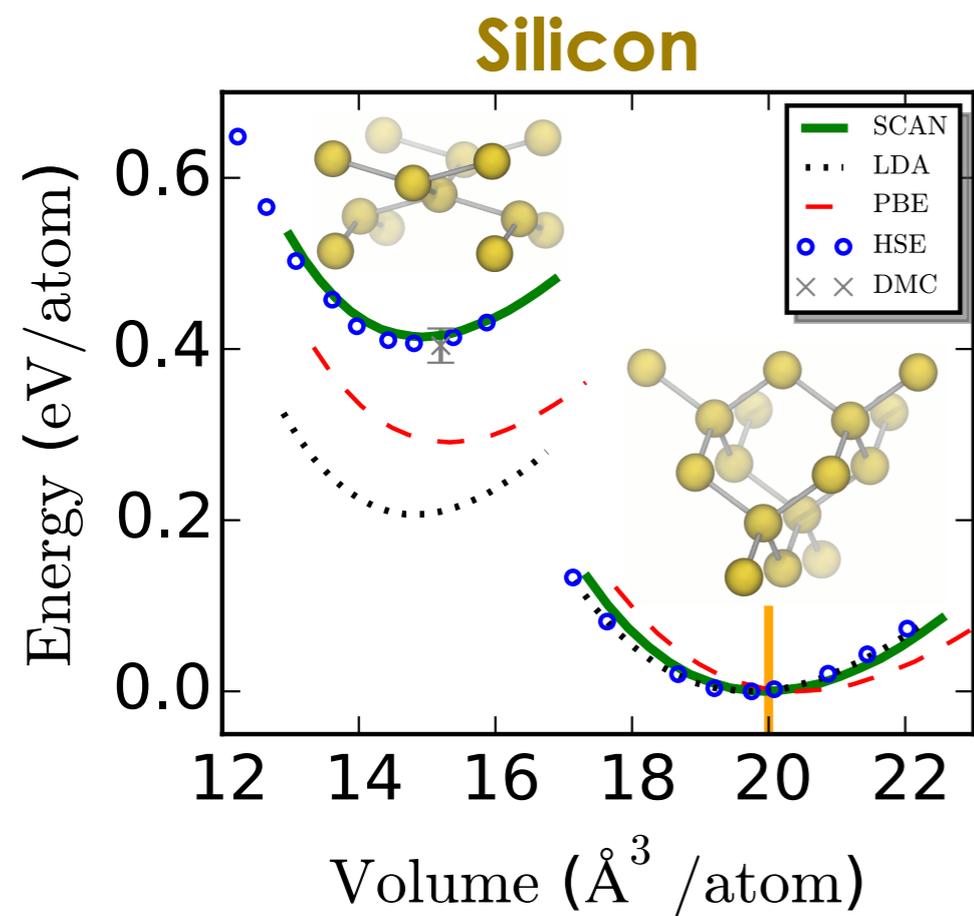
K_0 = bulk modulus

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- Equation of State & Bulk Moduli

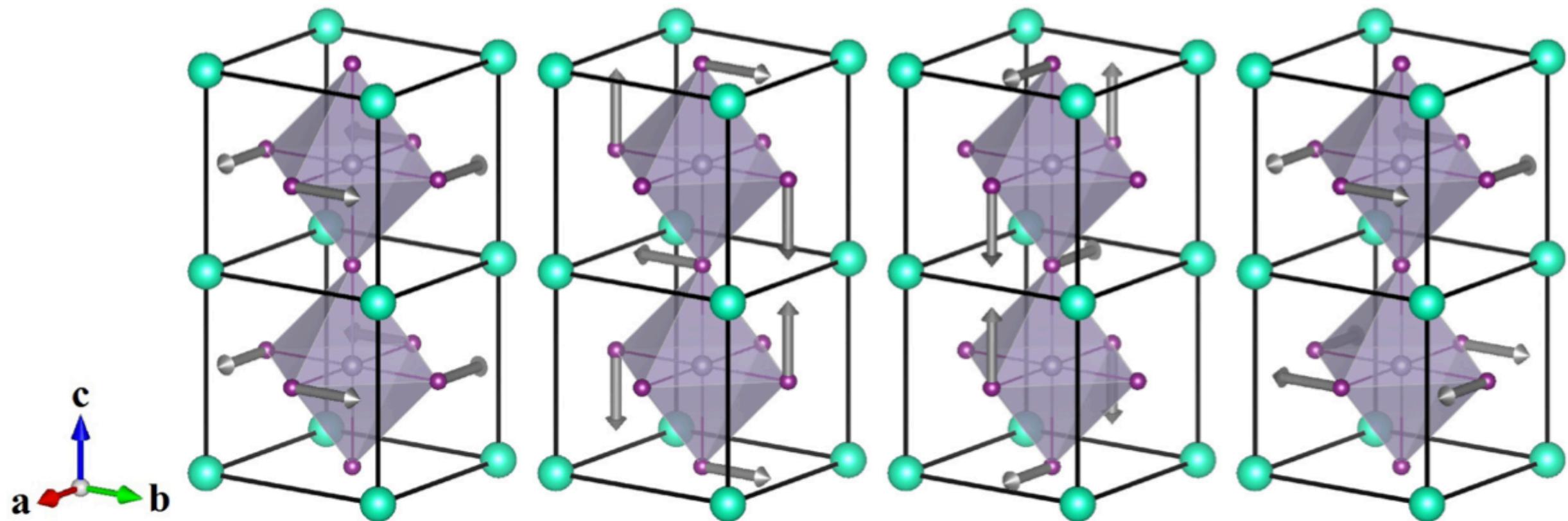


Phase Stability & Defect Formation



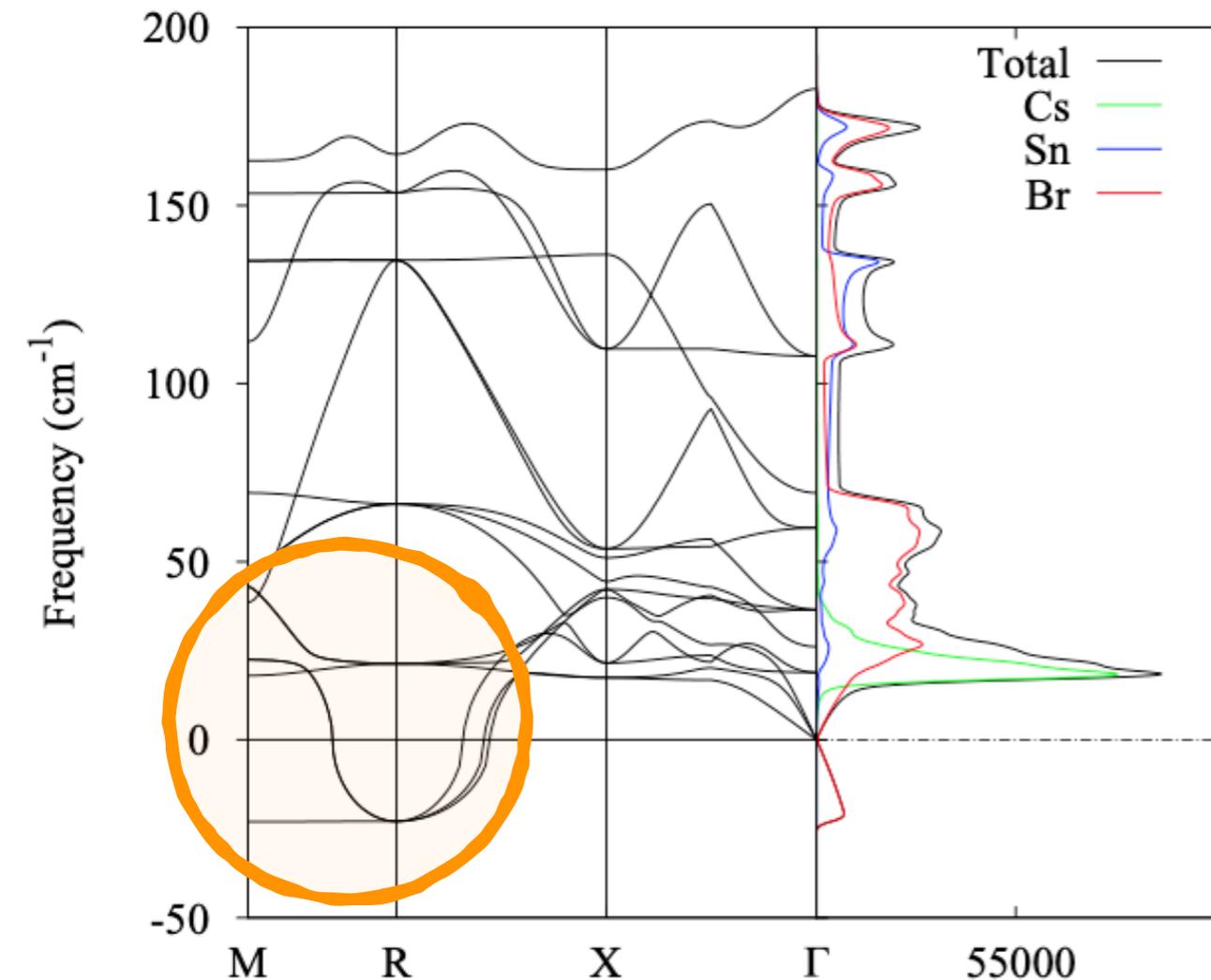
Why care about Forces?

- Forces (derivatives of total energy) tell us how **atoms move**
- Structural relaxation, cell optimization (stresses), mechanics
- **Molecular dynamics** → fluctuations, vibrations, diffusion, reactions
- **Let's look at vibrations (phonons) in a model material, CsSnBr₃**

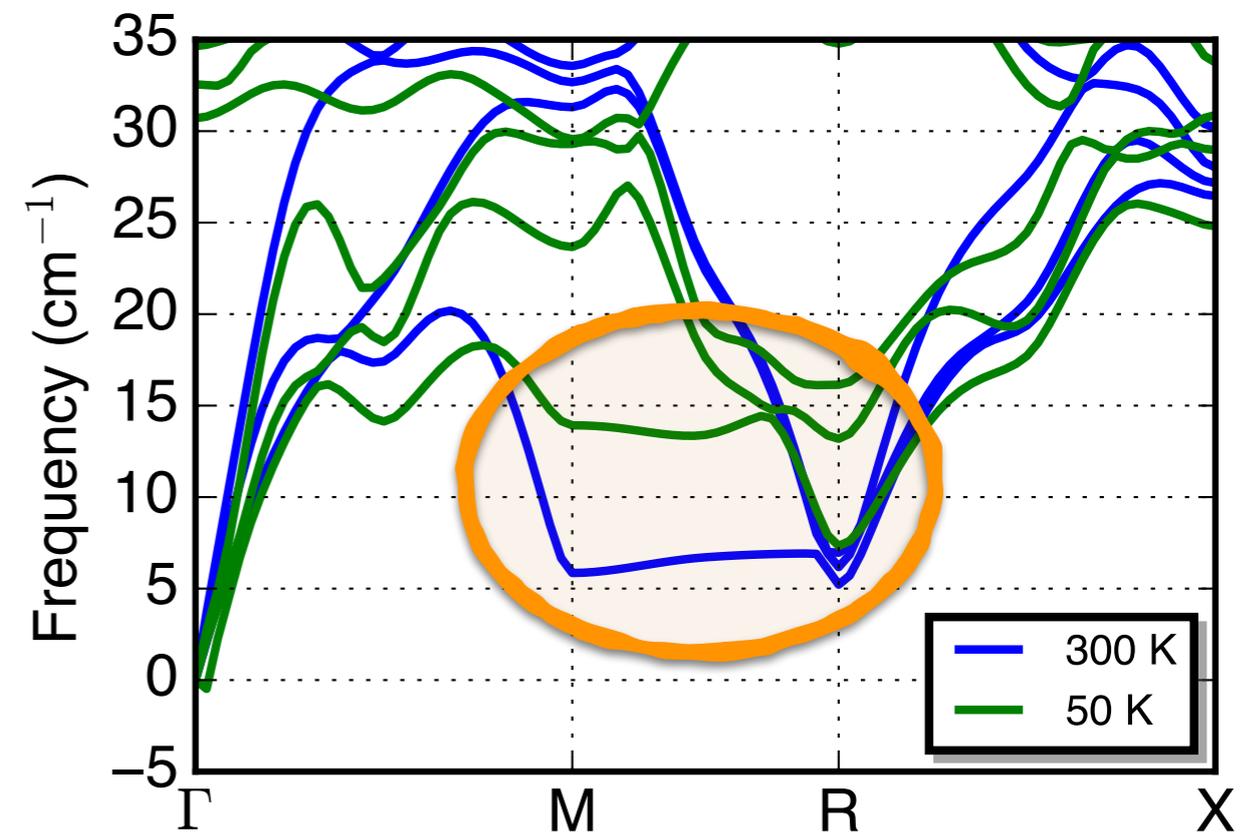


Phonons from harmonic approximations vs. MD simulations

- Phonon dispersion curves for CsSnBr₃

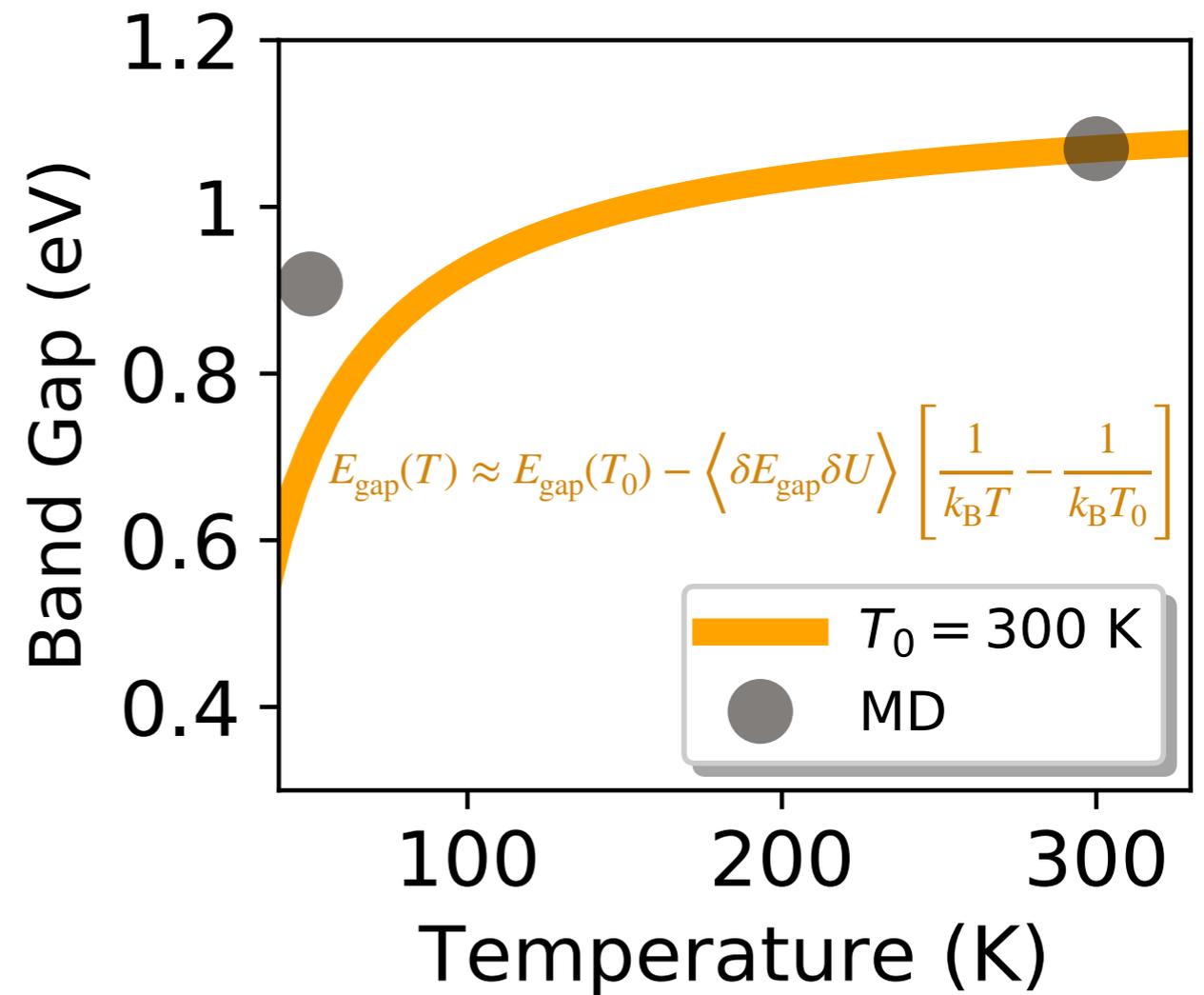
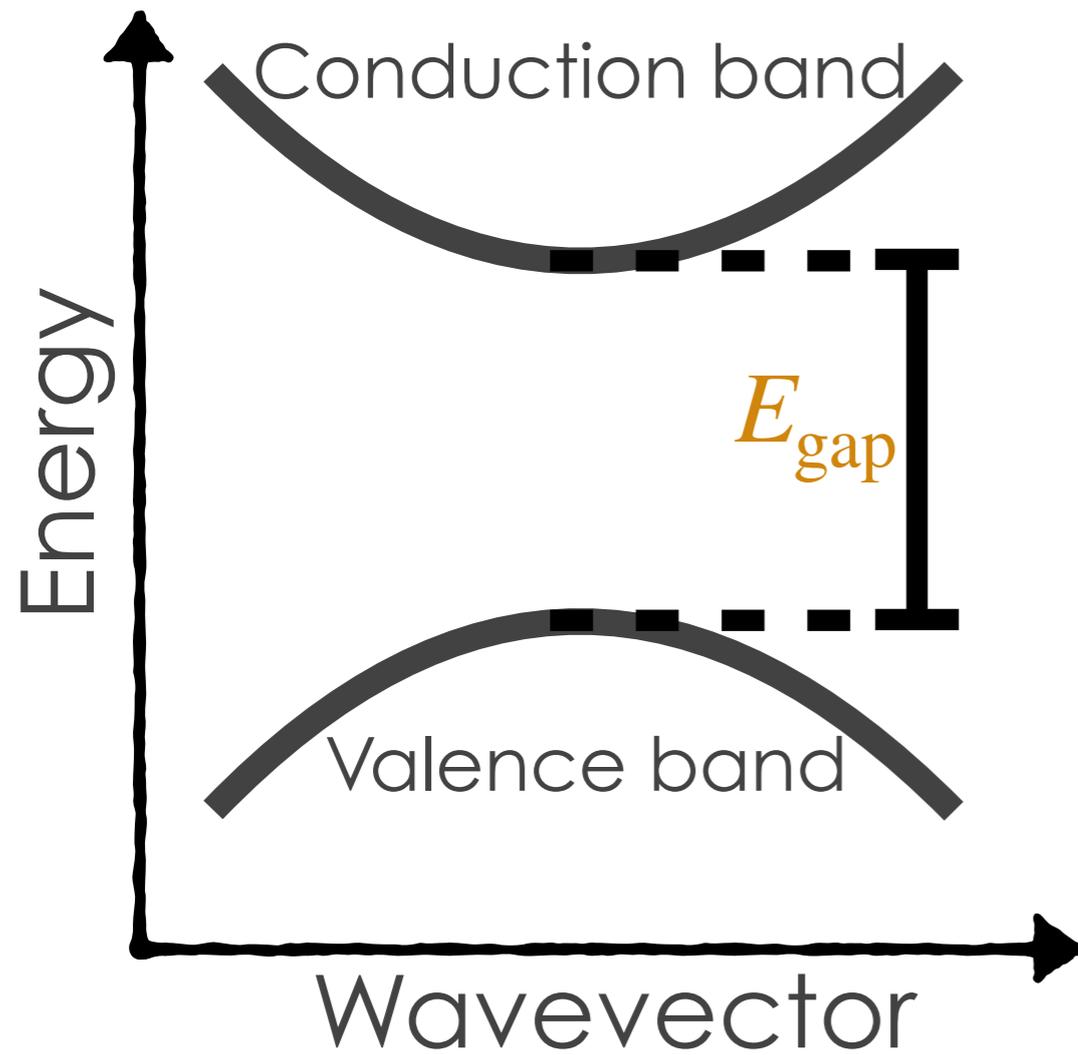


- Negative frequencies!**
= unstable / soft modes



- Anharmonic effects & thermal renormalization included in MD sims

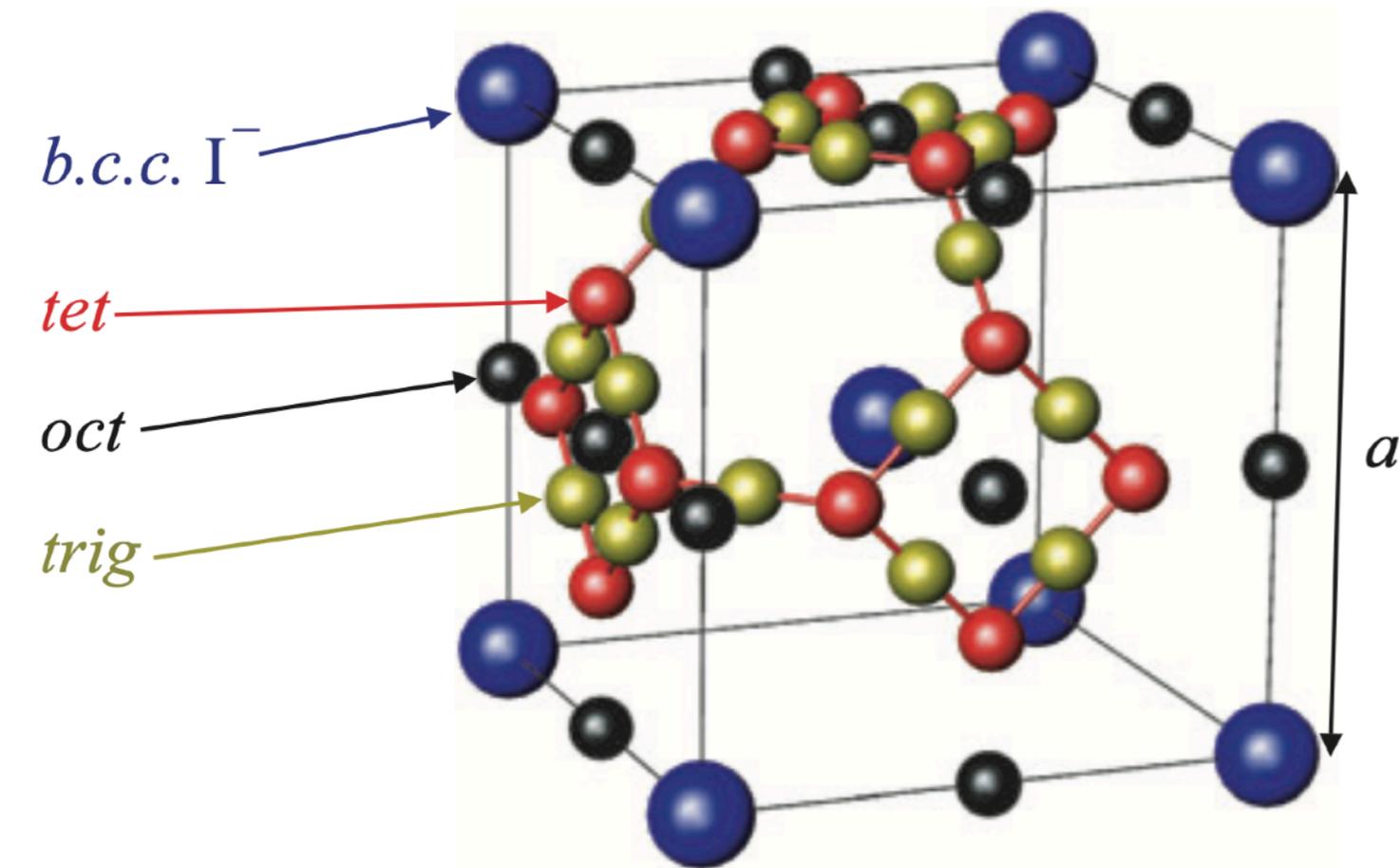
Band gap predictions must include anharmonicity



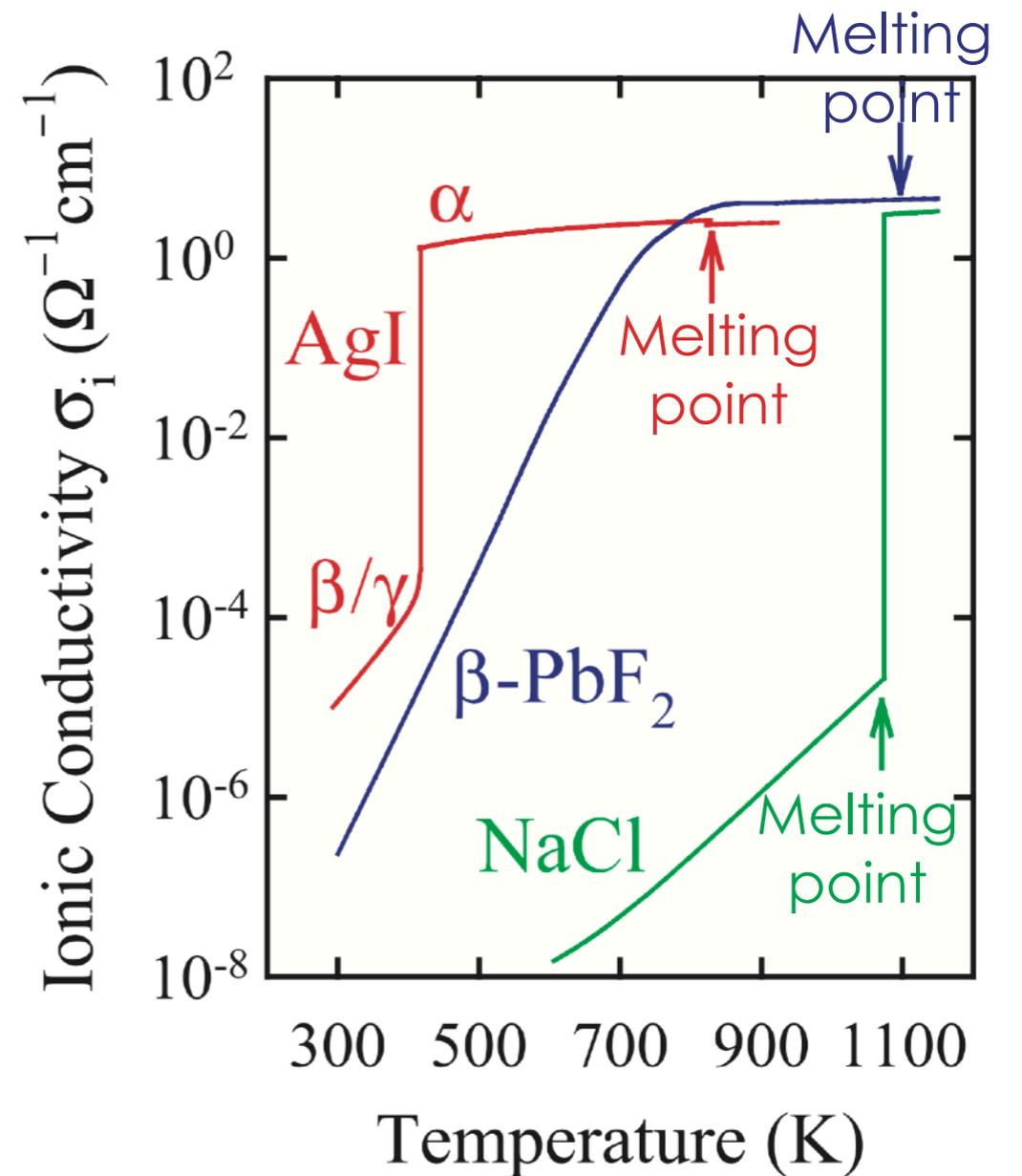
- Cannot describe band gap of anharmonic materials with linear/harmonic theories
- **Need anharmonicity** (from *hidden electronic disorder*)

Cation diffusion in Silver Iodide

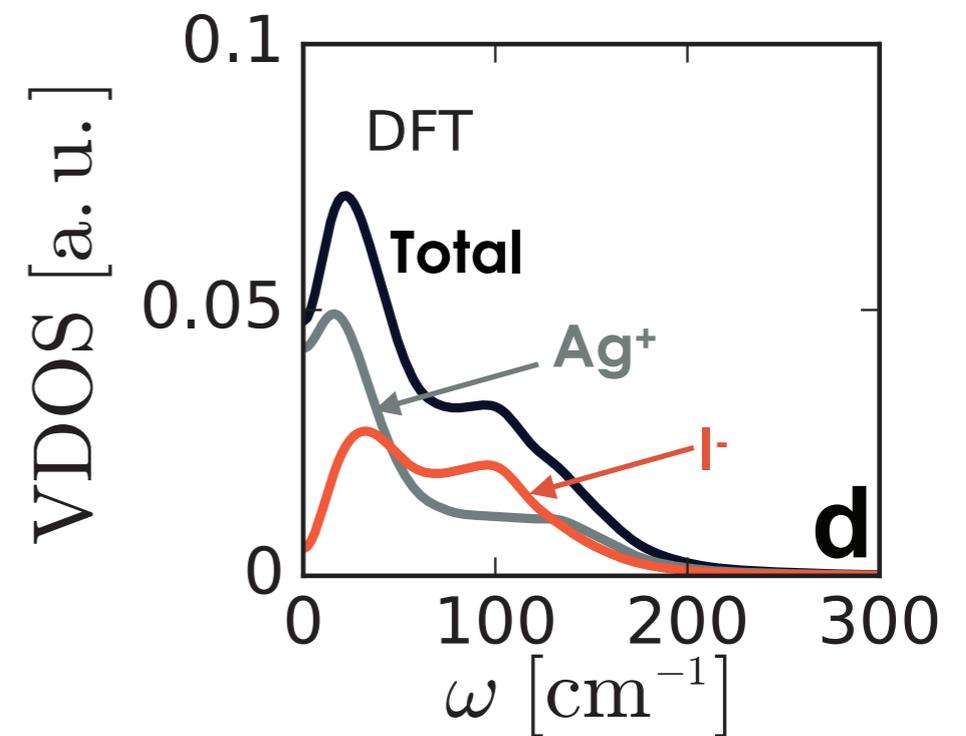
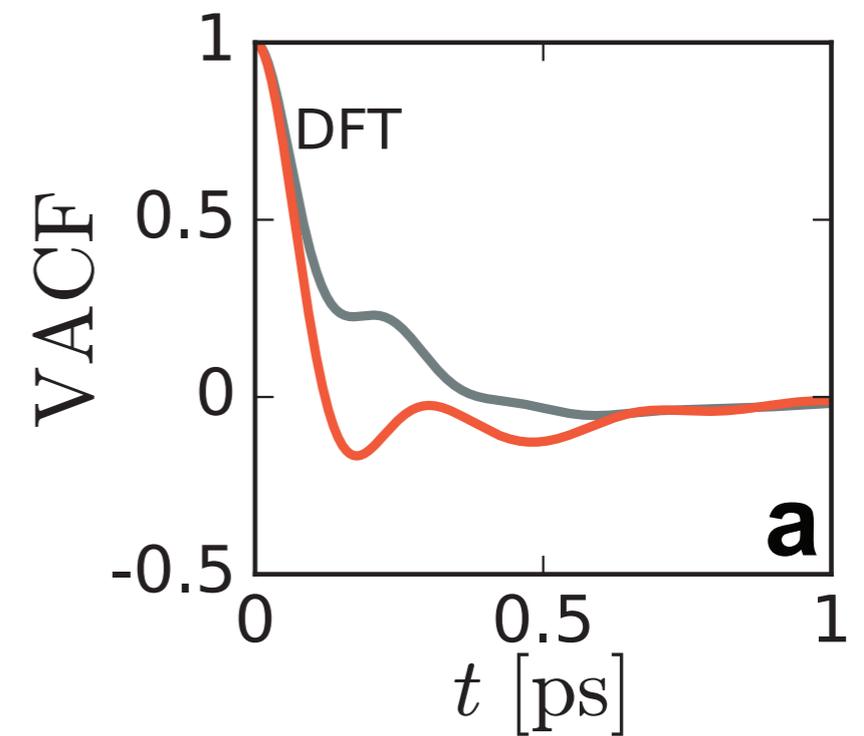
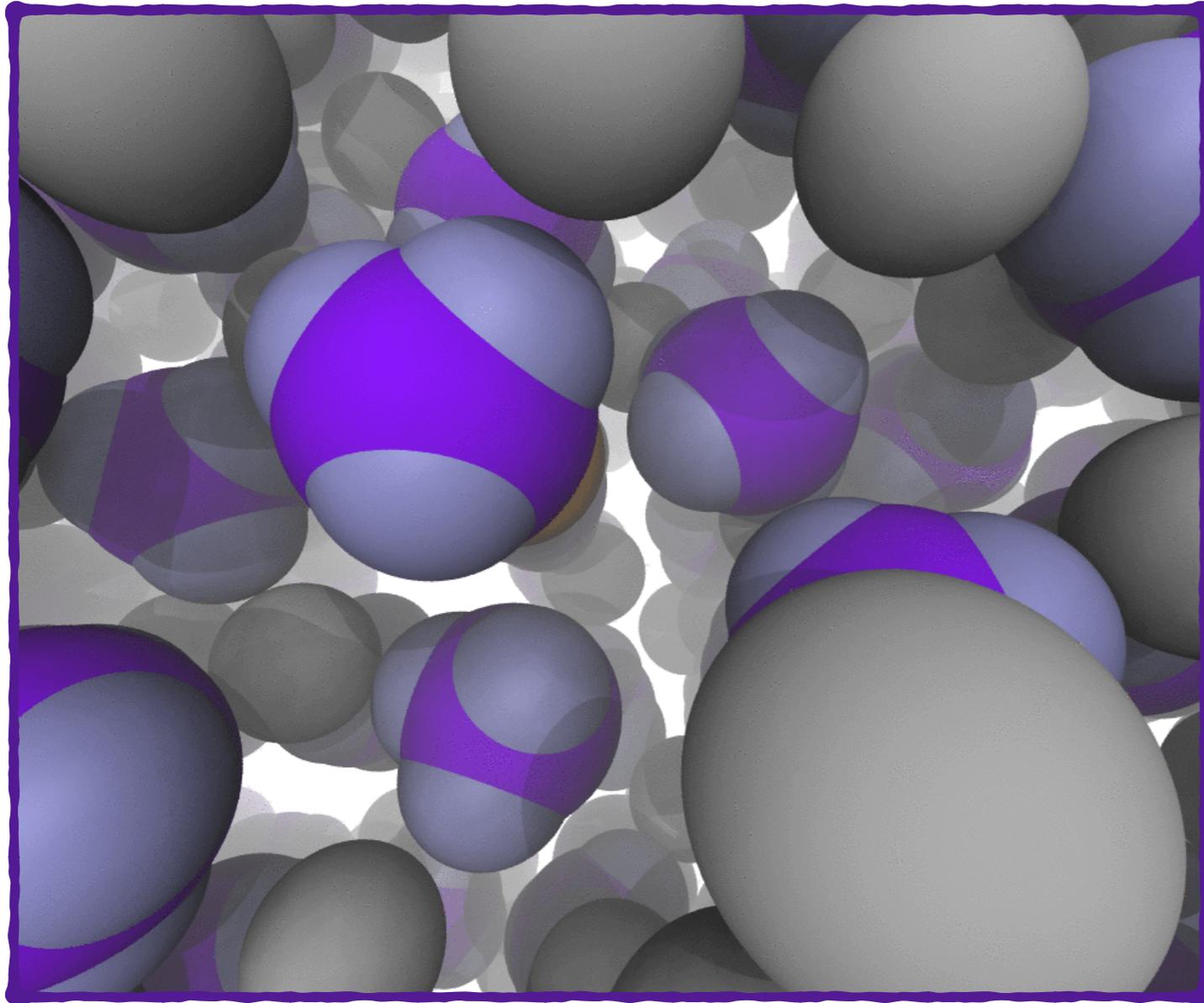
- Iodide sits on cubic lattice, Ag^+ diffuses



S. Hull, *Rep Prog Phys* 2004



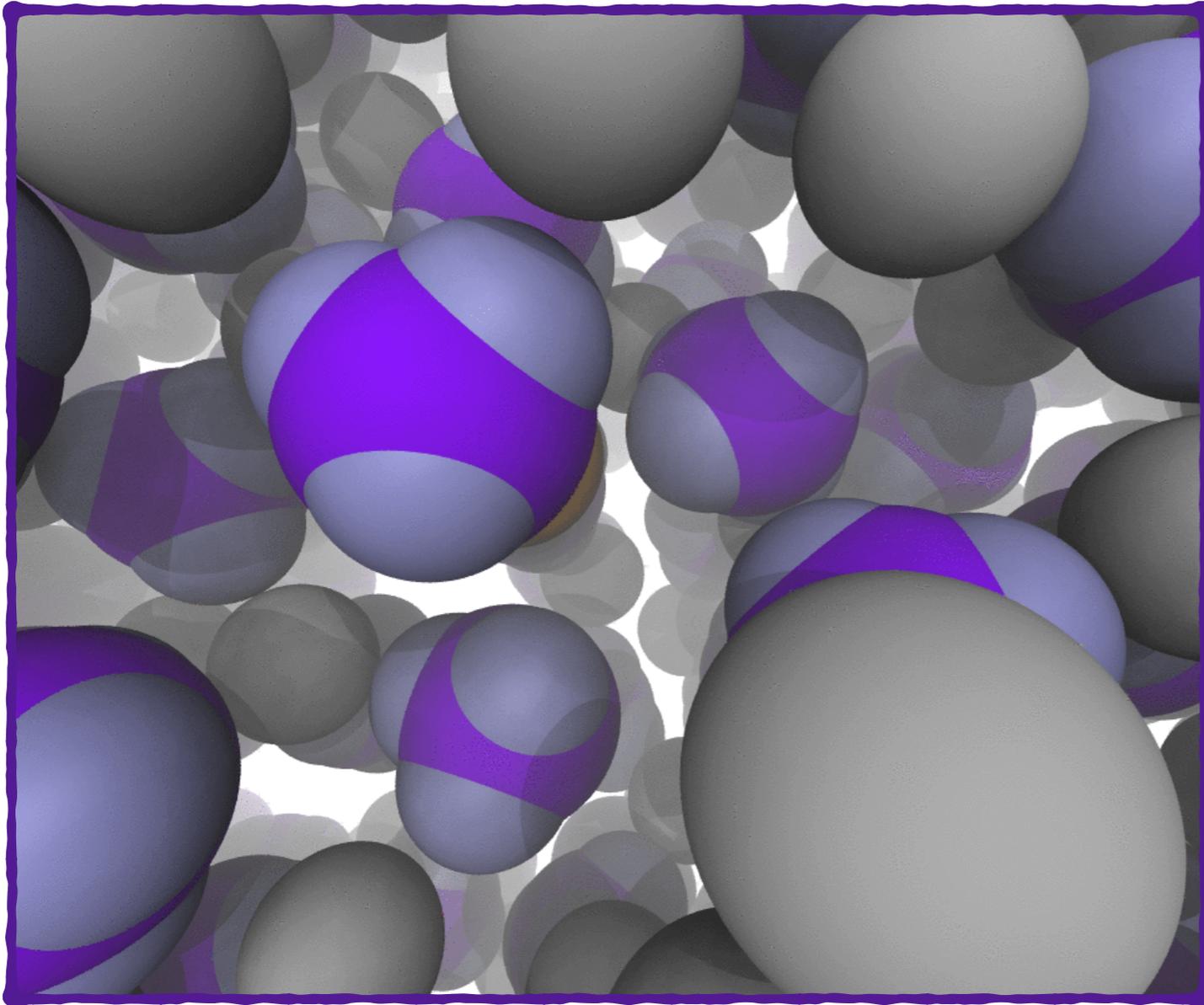
Cation diffusion in Silver Iodide



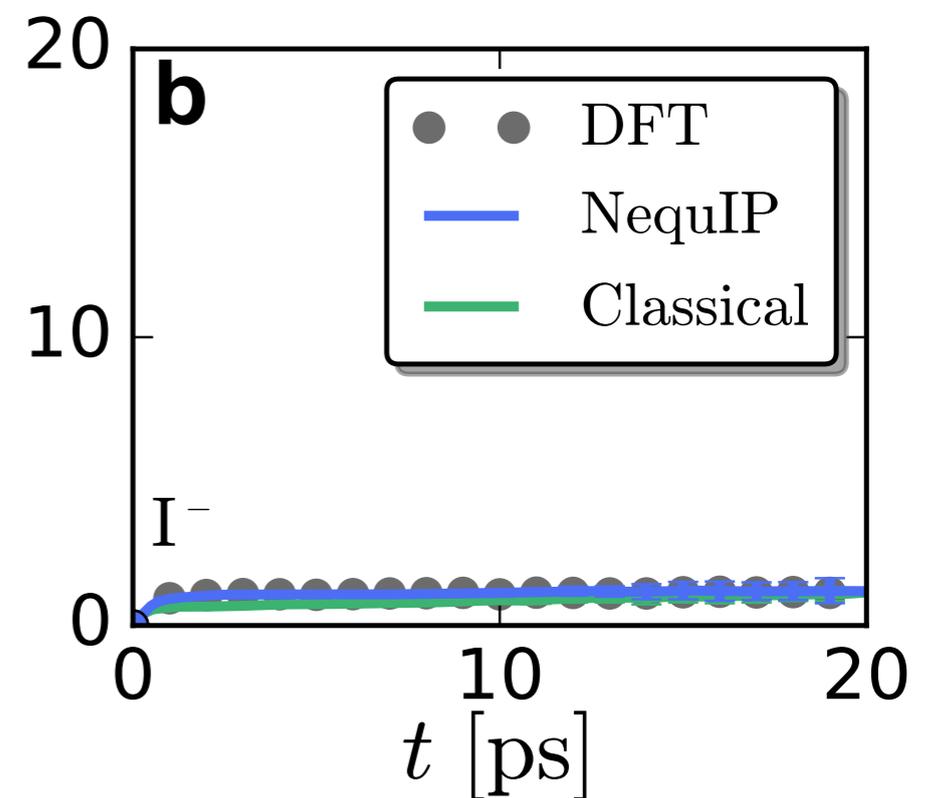
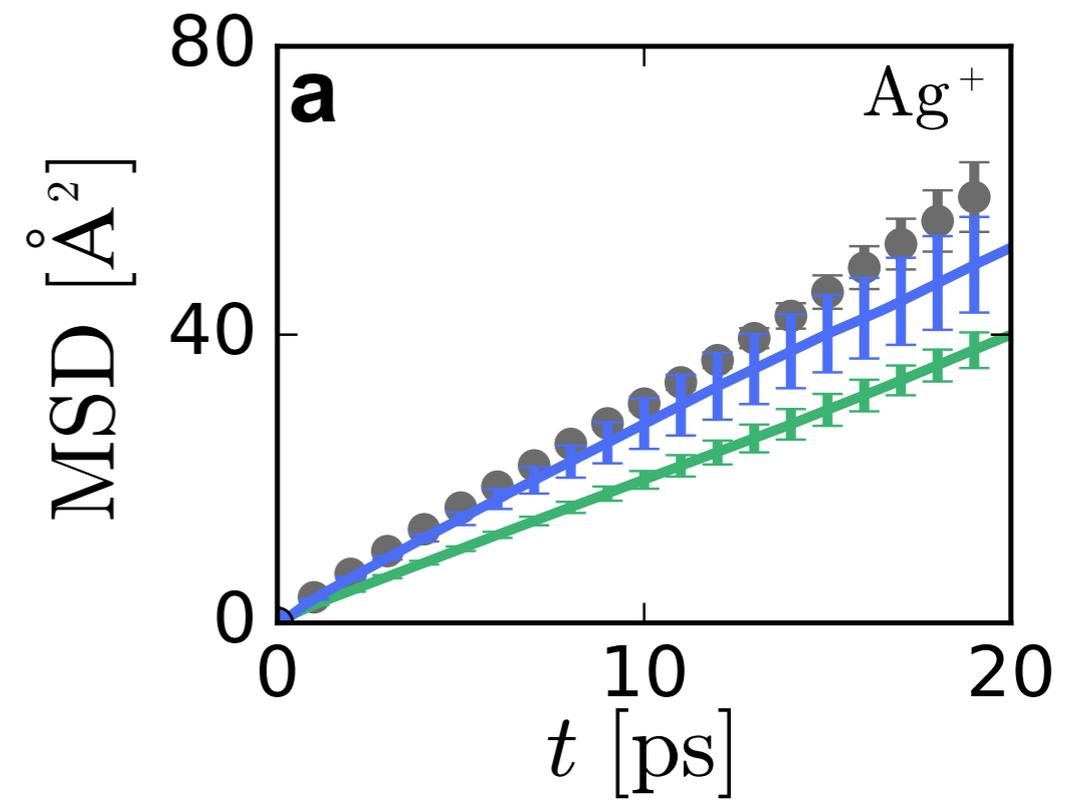
- $VACF(t) = \langle \mathbf{v}(t) \cdot \mathbf{v}(0) \rangle$

- $VDOS(\omega) = \int_{-\infty}^{\infty} \langle \mathbf{v}(t) \cdot \mathbf{v}(0) \rangle e^{-i\omega t} dt ; VDOS(\omega = 0) = 3D$

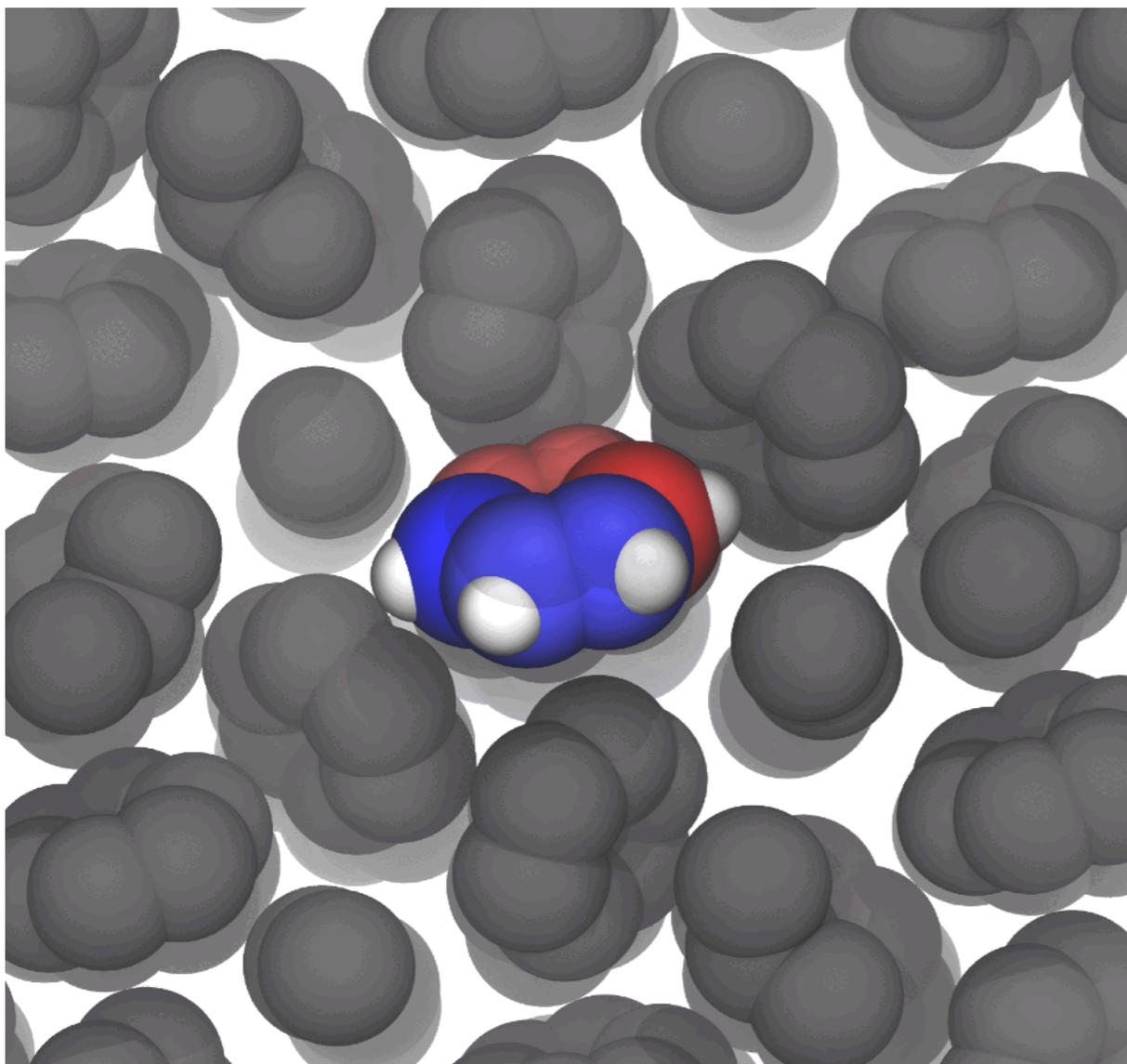
Cation diffusion in Silver Iodide



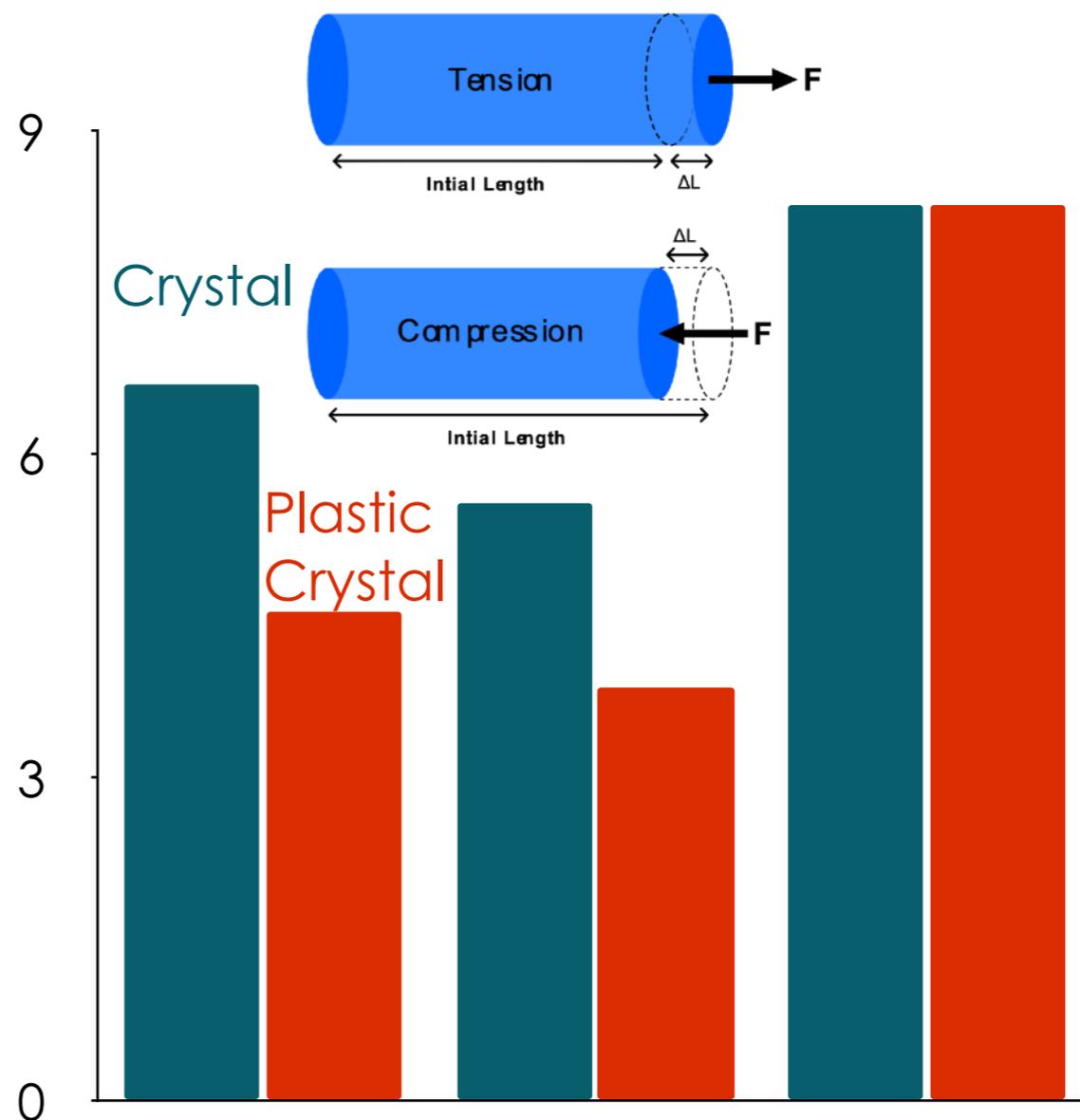
- $\text{MSD}(t) = \left\langle \left| \mathbf{r}(t) - \mathbf{r}(0) \right|^2 \right\rangle$
- $\lim_{t \rightarrow \infty} \text{MSD}(t) = 6Dt$



Rotational Disorder & Mechanics

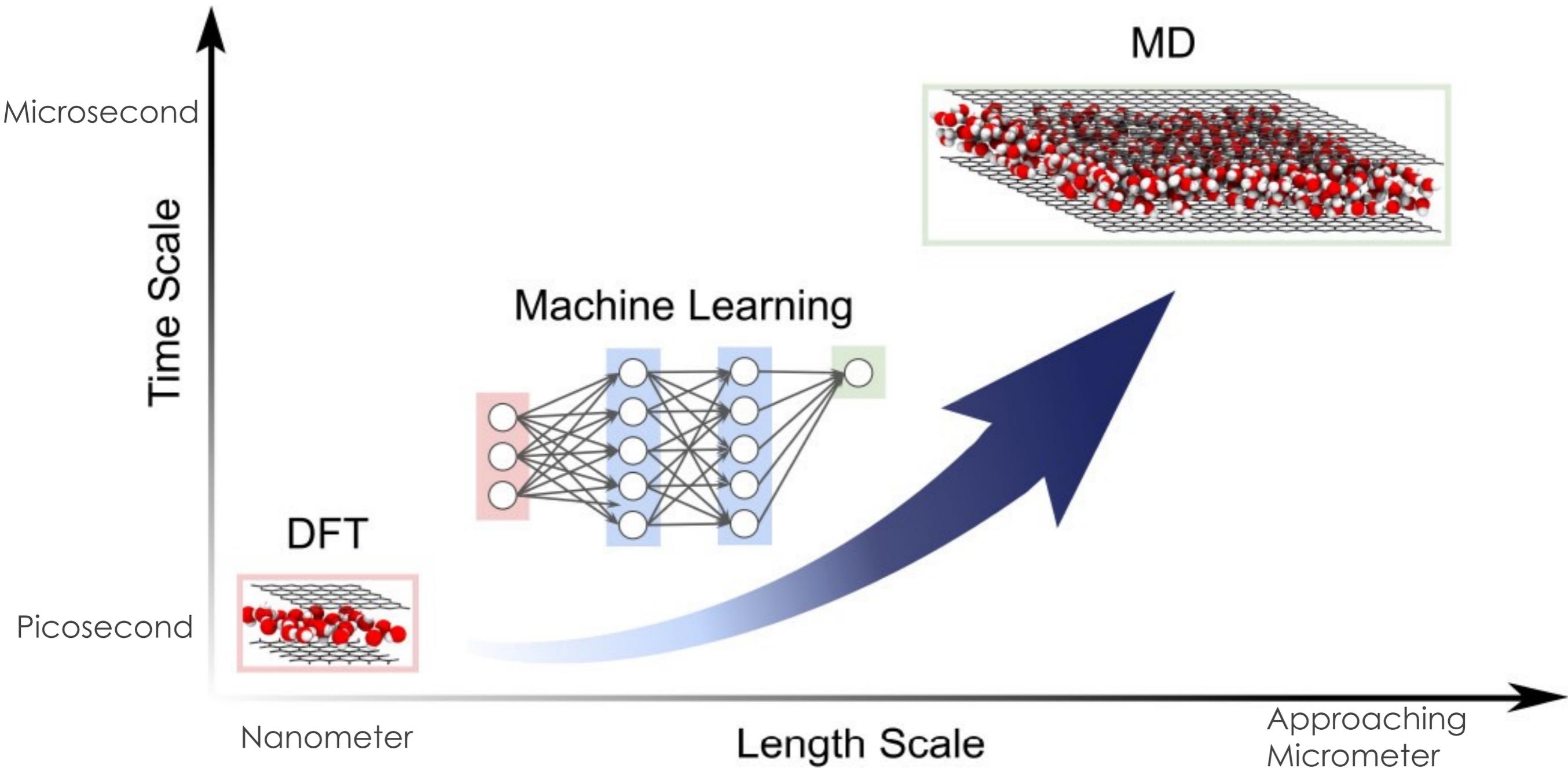


- **Plastic Crystals** have translational order & rotational disorder



- Rotational disorder can weaken mechanical properties (e.g. Young's moduli)

Bridging the gap with machine learning

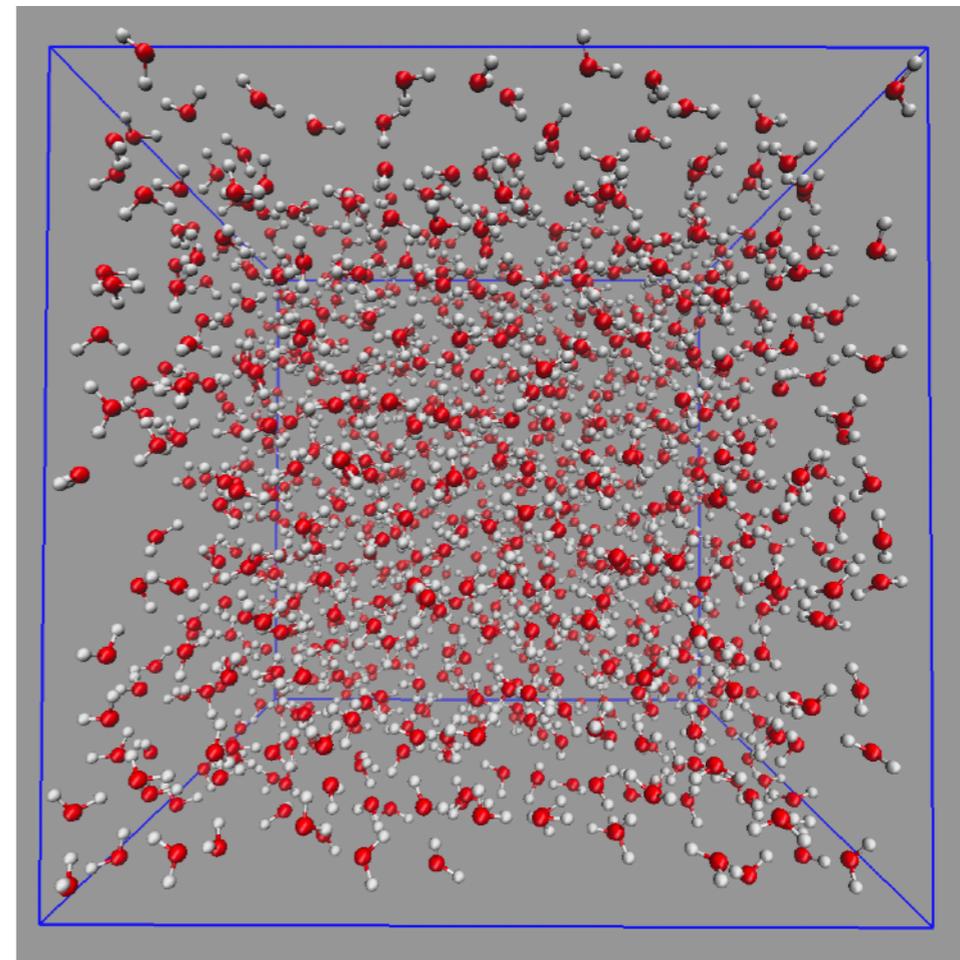
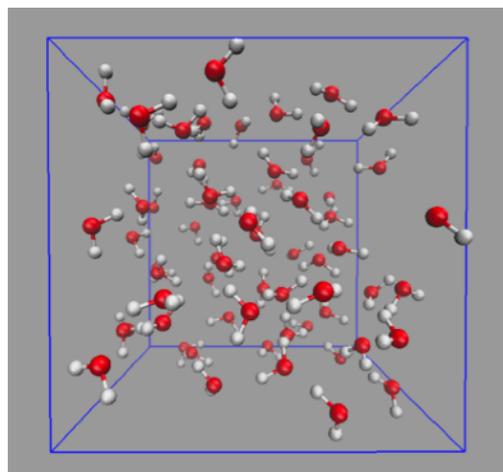
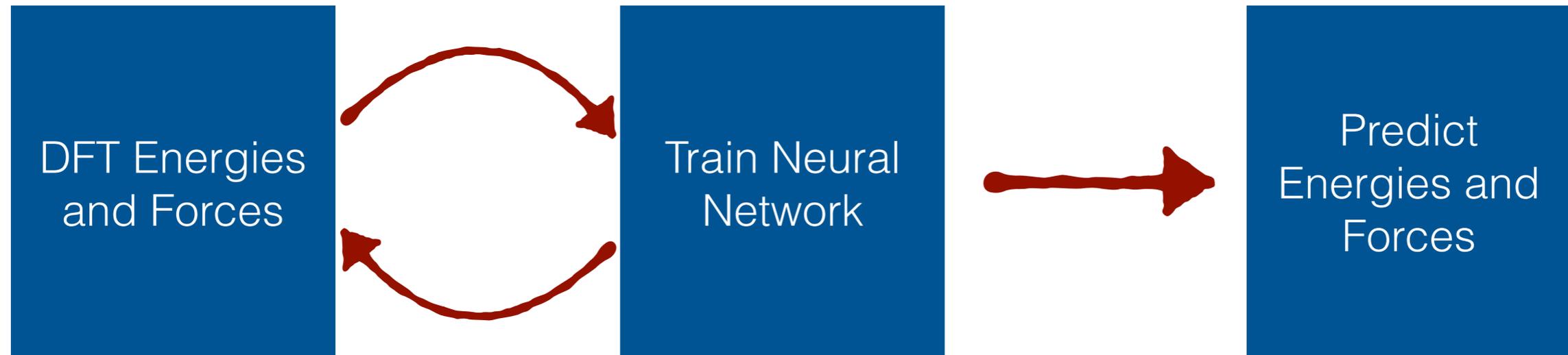


- **Machine learning moves us in the right direction**

- *Still developing techniques that include important physics...*

Use machine learning to model interatomic interactions

- Machine learning to make **neural network potentials**



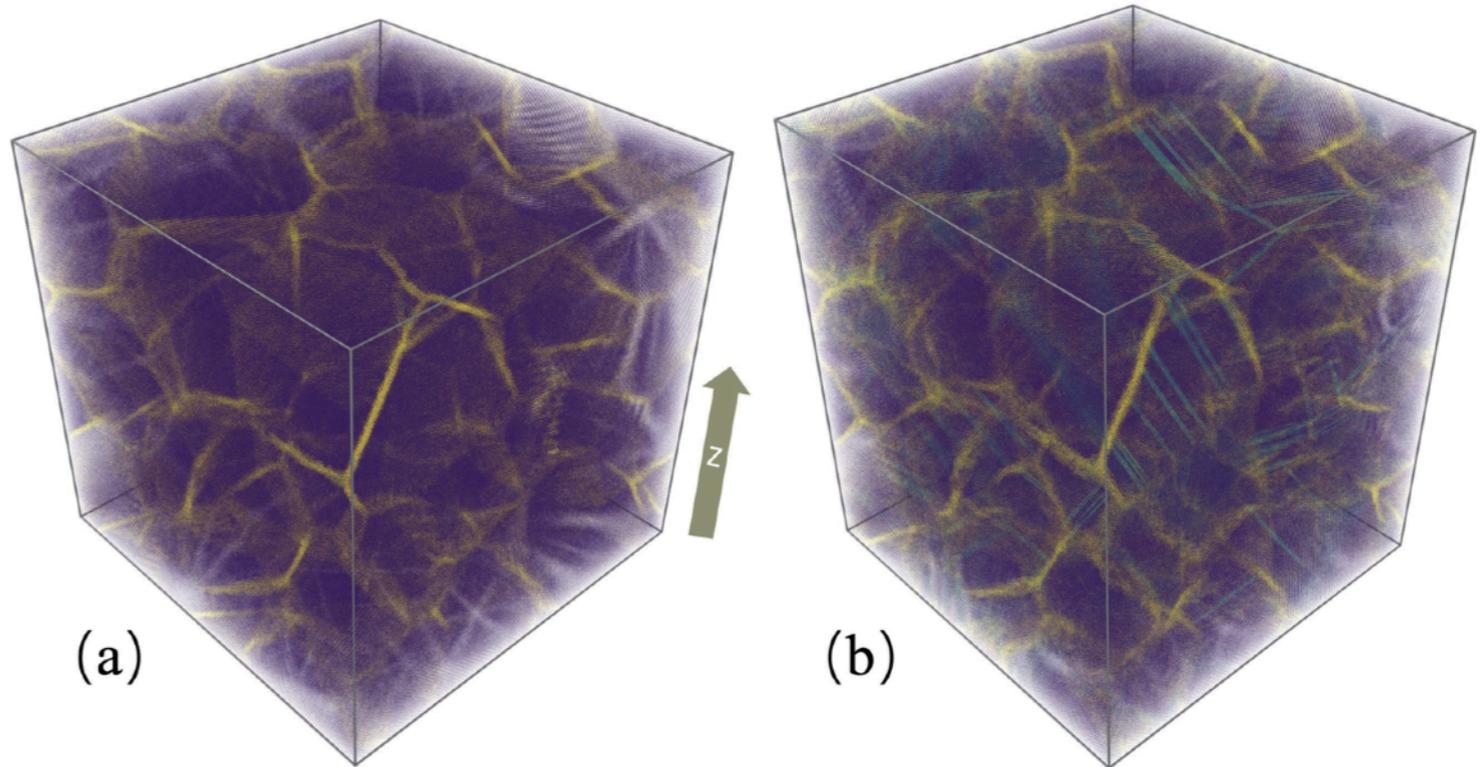
- Can model large systems with DFT accuracy!

Scalability enables increase in scales reachable in simulations

- Can model huge systems

Pushing the Limit of Molecular Dynamics with *Ab Initio* Accuracy to 100 Million Atoms with Machine Learning

Weile Jia*, Han Wang†, Mohan Chen‡, Denghui Lu‡, Lin Lin*¶, Roberto Car||, Weinan E||, Linfeng Zhang|| §
 *University of California, Berkeley, Berkeley, USA
 Email: jiaweile@berkeley.edu, linlin@math.berkeley.edu
 †Laboratory of Computational Physics, Institute of Applied Physics and Computational Mathematics, Beijing, China
 Email: wang_han@iapcm.ac.cn
 ‡CAPT, HEDPS, College of Engineering, Peking University, Beijing, China
 Email: mohanchen@pku.edu.cn, denghuilu@pku.edu.cn
 ¶Lawrence Berkeley National Laboratory, Berkeley, USA
 || Princeton University, Princeton, USA

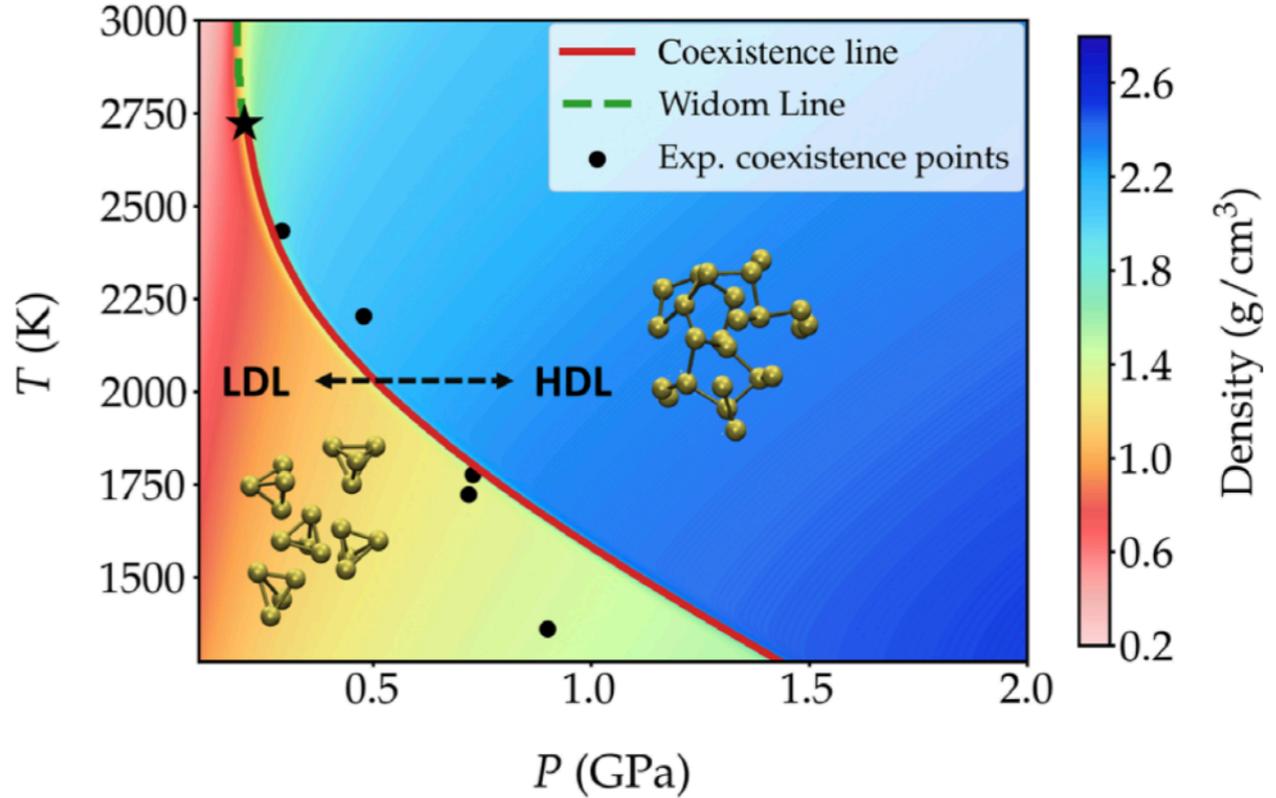


- Can model long timescales for thermodynamics & kinetics

PHYSICAL REVIEW LETTERS 127, 080603 (2021)

Liquid-Liquid Critical Point in Phosphorus

Manyi Yang, Tarak Karmakar, and Michele Parrinello
 Italian Institute of Technology, Via Melen 83, 16152 Genova, Italy



- All at ab initio accuracy!

**To the board to discuss
the theory behind these calculations**

What are Molecular Dynamics (MD) Simulations?

- ❖ Numerical solution of classical equations of motion (Newton's Equations) for a molecular system

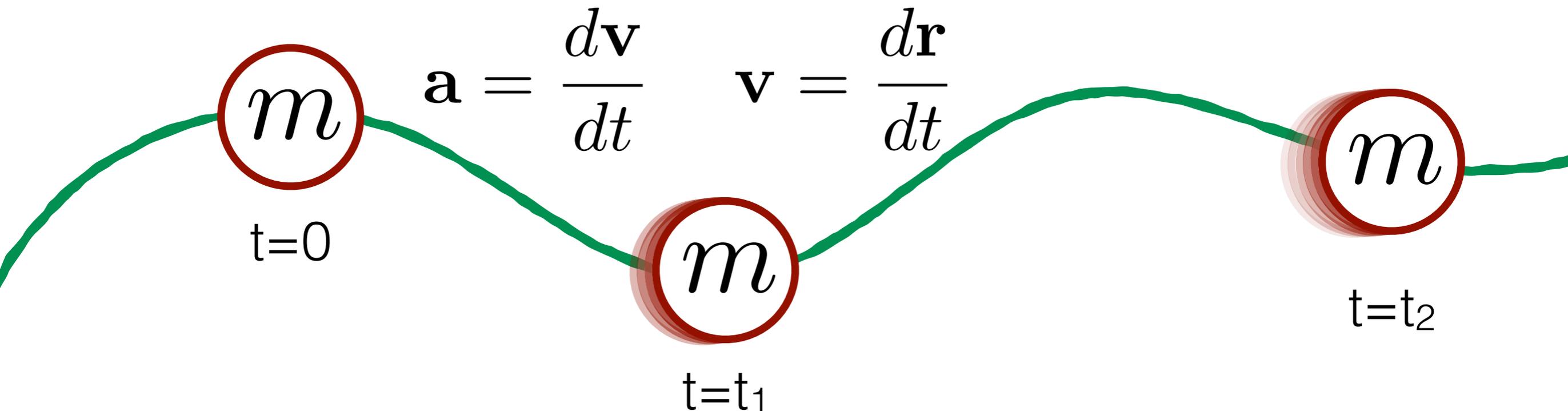
$$\mathbf{F} = m\mathbf{a}$$

What are Molecular Dynamics (MD) Simulations?

- ❖ Numerical solution of classical equations of motion (Newton's Equations) for a molecular system

$$\mathbf{F} = m\mathbf{a}$$

- ❖ Given an initial condition (positions and momenta), we can numerically evolve the system in time

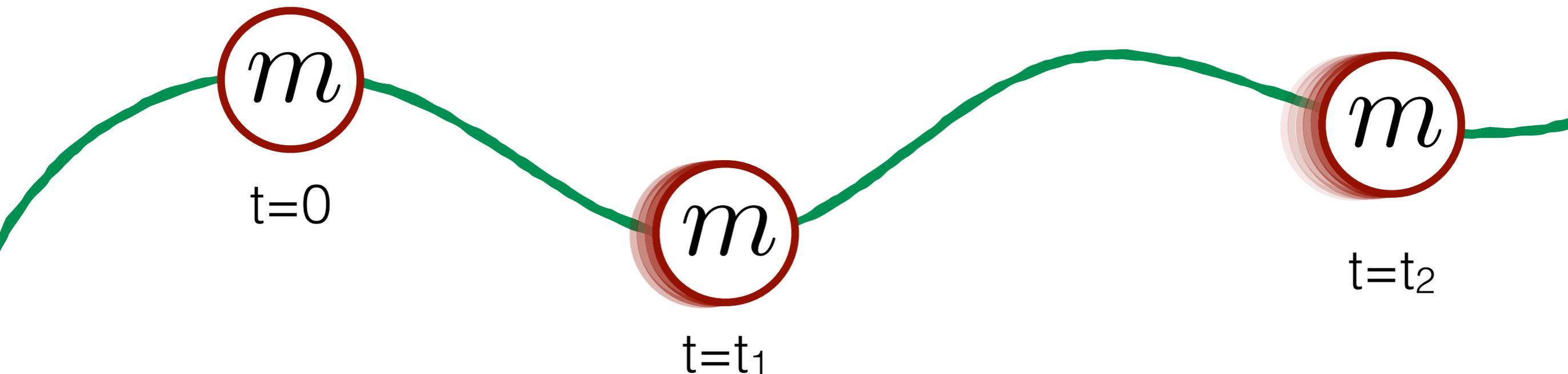


Integrating the equations of motion

- ✦ Velocity Verlet algorithm (Taylor expansion)

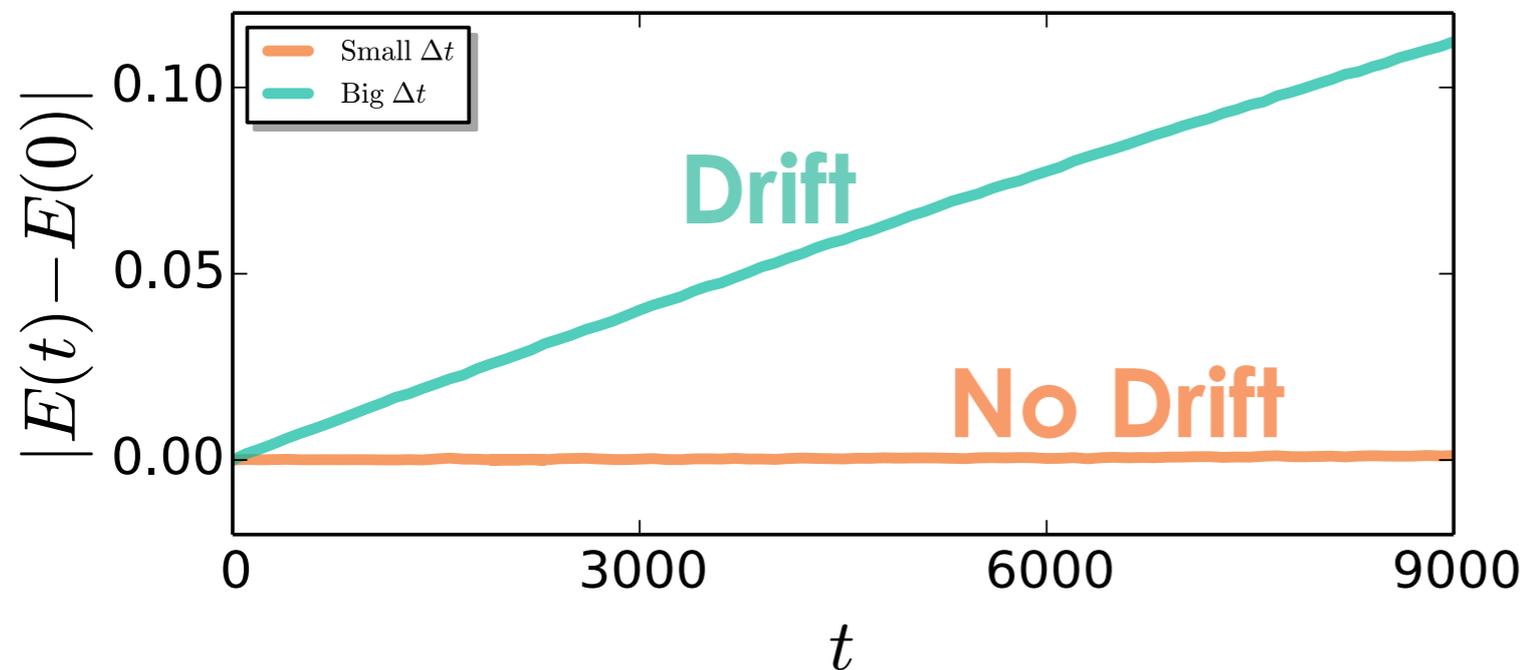
$$\mathbf{r}(t + \Delta t) = \mathbf{r}(t) + \mathbf{v}(t)\Delta t + \frac{1}{2}\mathbf{a}(t)\Delta t^2$$

$$\mathbf{v}(t + \Delta t) = \mathbf{v}(t) + \Delta t \frac{\mathbf{a}(t) + \mathbf{a}(t + \Delta t)}{2}$$



How to choose the timestep?

- ❖ Want Δt large enough to simulate for long times
- ❖ Need Δt small enough to accurately integrate EOM; error $O(\Delta t^2)$



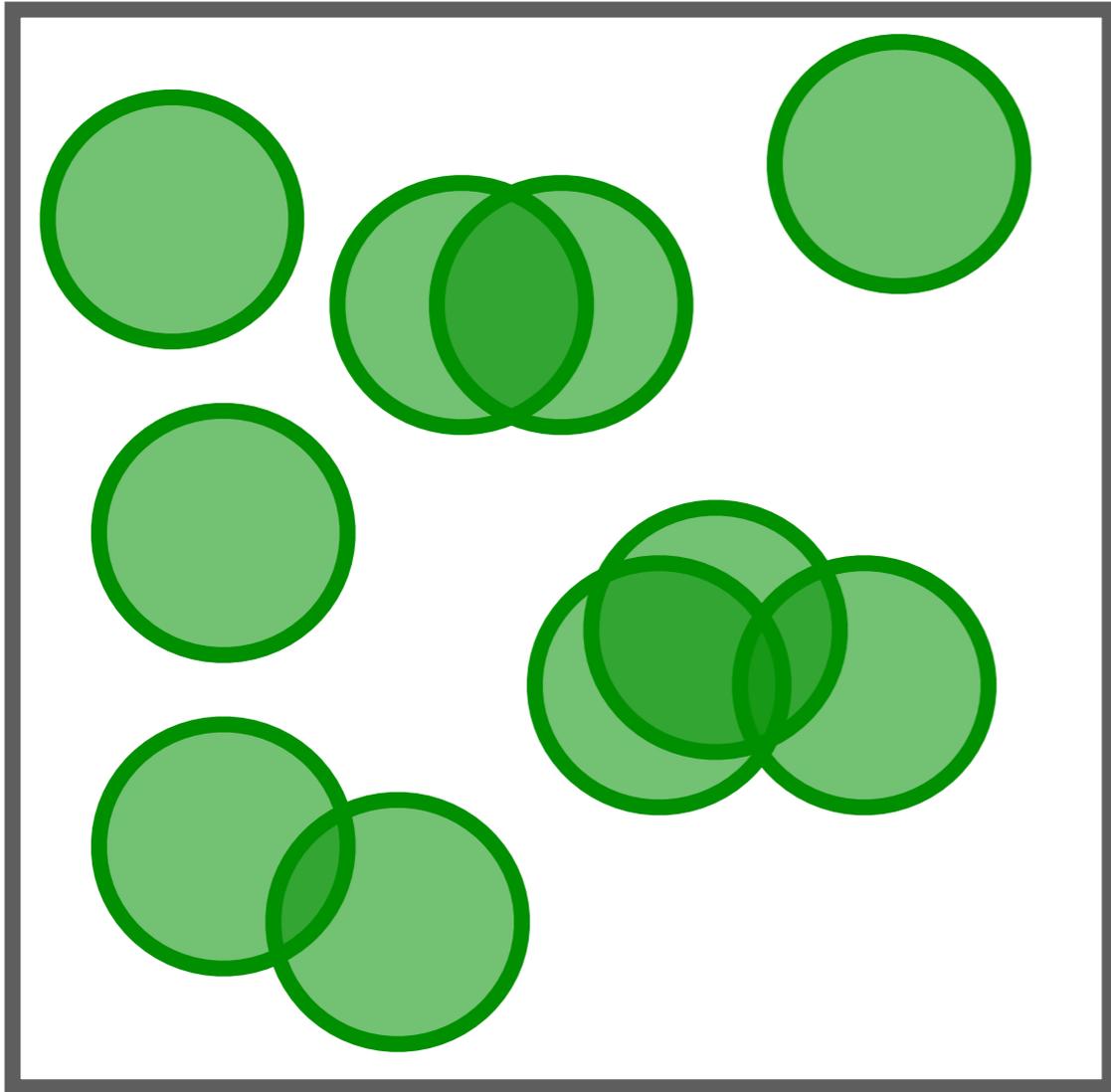
- ❖ Δt should conserve constant of motion; here, the energy E

- ❖ Typically Δt smaller than fastest timescale in the system

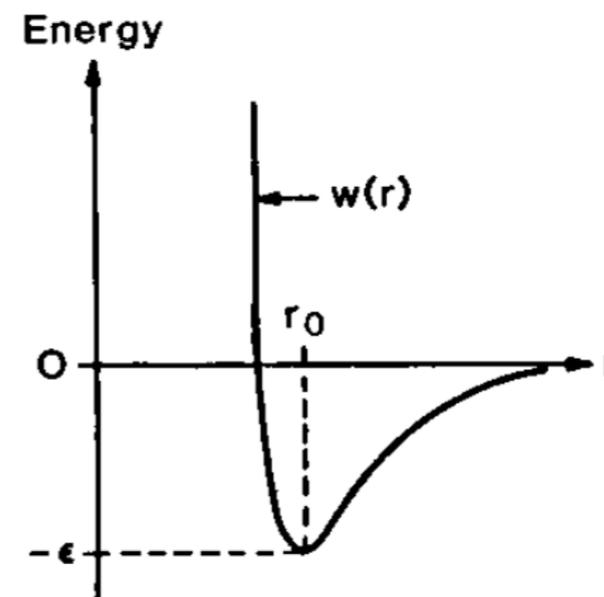
system (fastest motion)	time step (1 fs = 10^{-15} s)
molecules (bond vibrations)	0.5 – 1.0 fs
molecules, rigid bonds (angle bending)	2.0 fs
atoms (translation)	5 – 10 fs
Lennard-Jones system	0.001 (dimensionless units)

Choosing Initial Conditions

- ❖ Randomly insert **N** molecules into simulation box



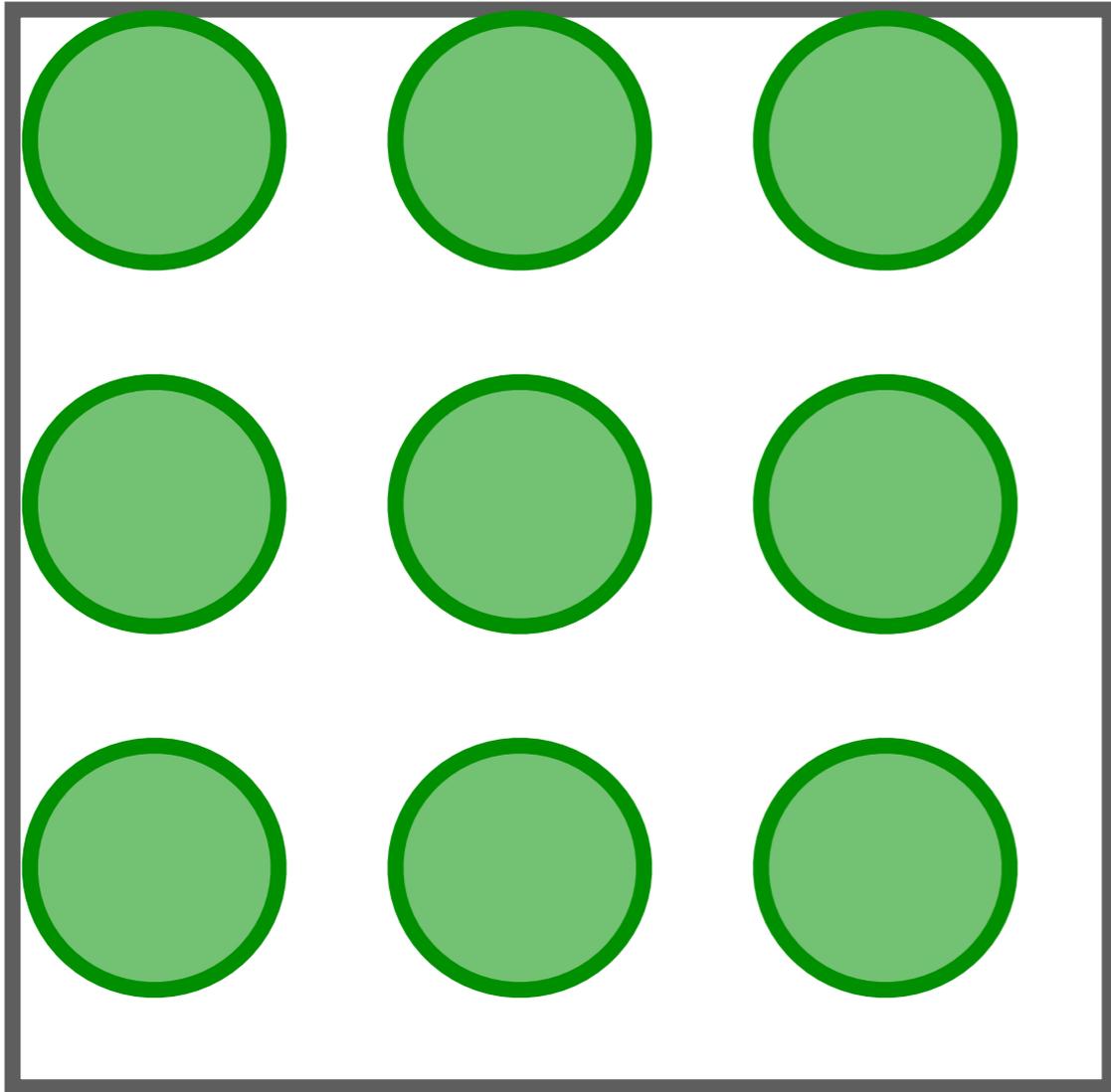
- ❖ Almost all intermolecular potentials include an excluded volume core (Pauli exclusion)



- ❖ Overlapping cores lead to high energy configurations
- ❖ Difficult to equilibrate: **Bad initial condition**
- ❖ Could make sure particles don't overlap...

Choosing Initial Conditions

- ❖ Put N molecules onto a lattice (often cubic)



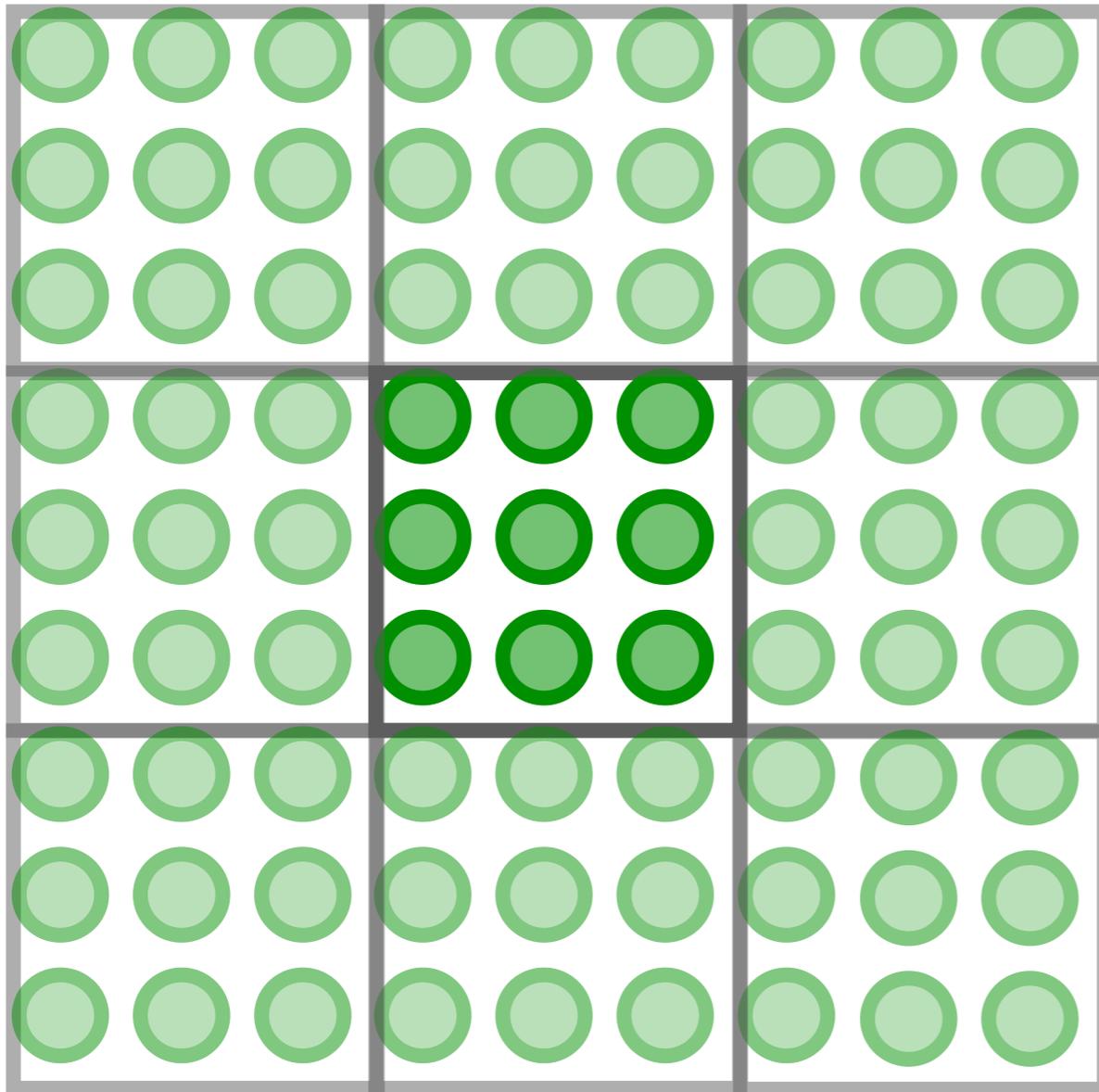
- ❖ Lattice will melt if it is not stable at this state point.
- ❖ Need to follow time evolution of “melting” to ensure you arrive at the desired state



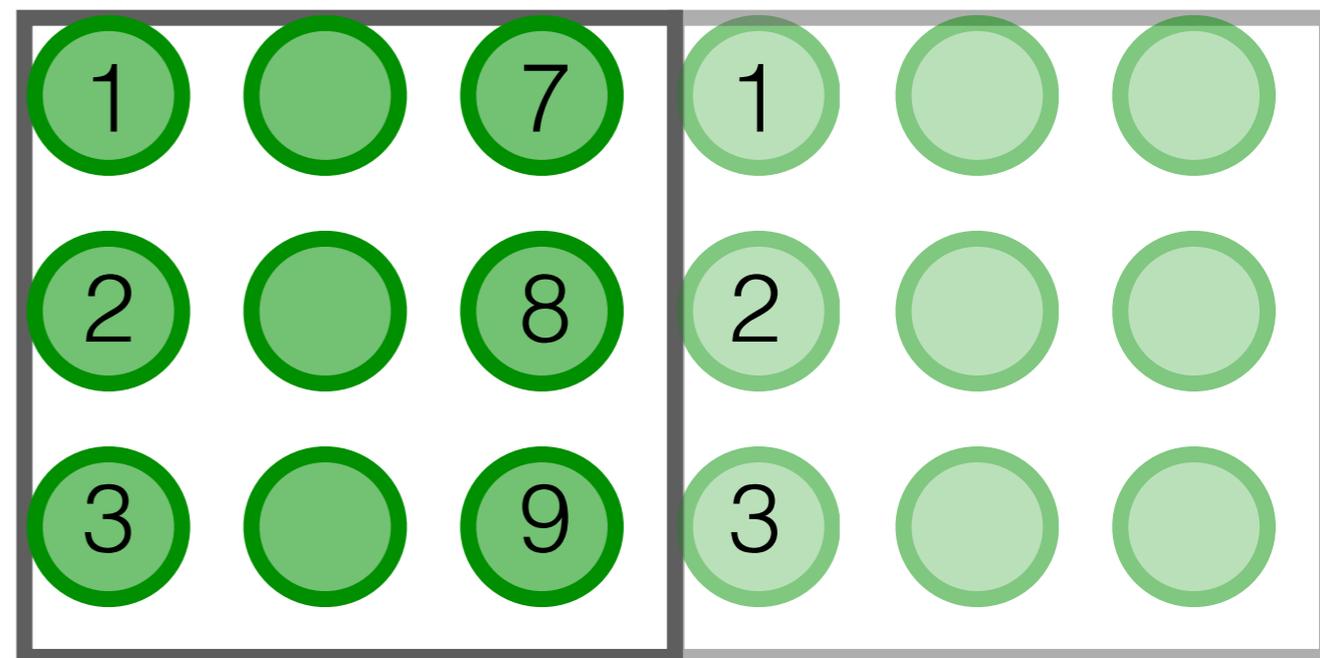
- ❖ *Why have I drawn the system off-center?*

Periodic Boundary Conditions

- ❖ Want to mimic an **infinite** system with **N** molecules



- ❖ Particles “feel” image of other particles outside on other side of boundary



- ❖ If particle “leaves” box, it “enters” on other side
- ❖ Be careful of artifacts! E.g. long-wavelength fluctuation suppression, finite-size effects,...

Initializing the Momenta

- ❖ Choose velocities randomly from a Gaussian distribution

$$P(v_x) = \left(\frac{m}{2\pi k_B T} \right)^{1/2} \exp \left(-\frac{mv_x^2}{2k_B T} \right)$$

- ❖ Need to correct so that there is no net momentum

$$\mathbf{P}_{\text{tot}} = \sum_{i=1}^N m_i \mathbf{v}_i = 0$$

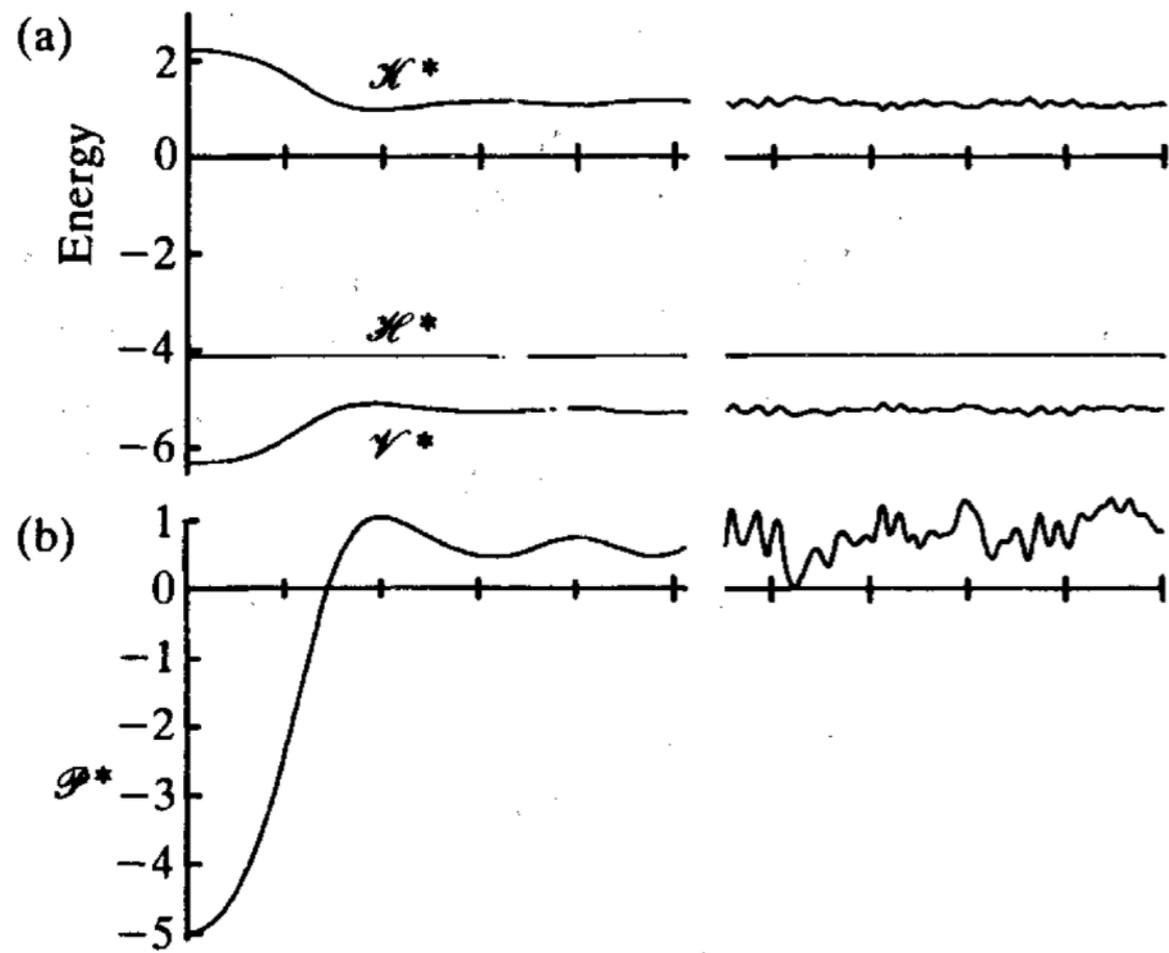
- ❖ System should rapidly equilibrate to Maxwell-Boltzmann distribution if initialized differently (e.g. uniform distribution)

Equilibrating the System

- ❖ Need to run for a period of time to allow the system to come to equilibrium at the state point of interest
- ❖ Monitor thermodynamic & structural quantities to ensure quantities no longer drift and oscillate about average values

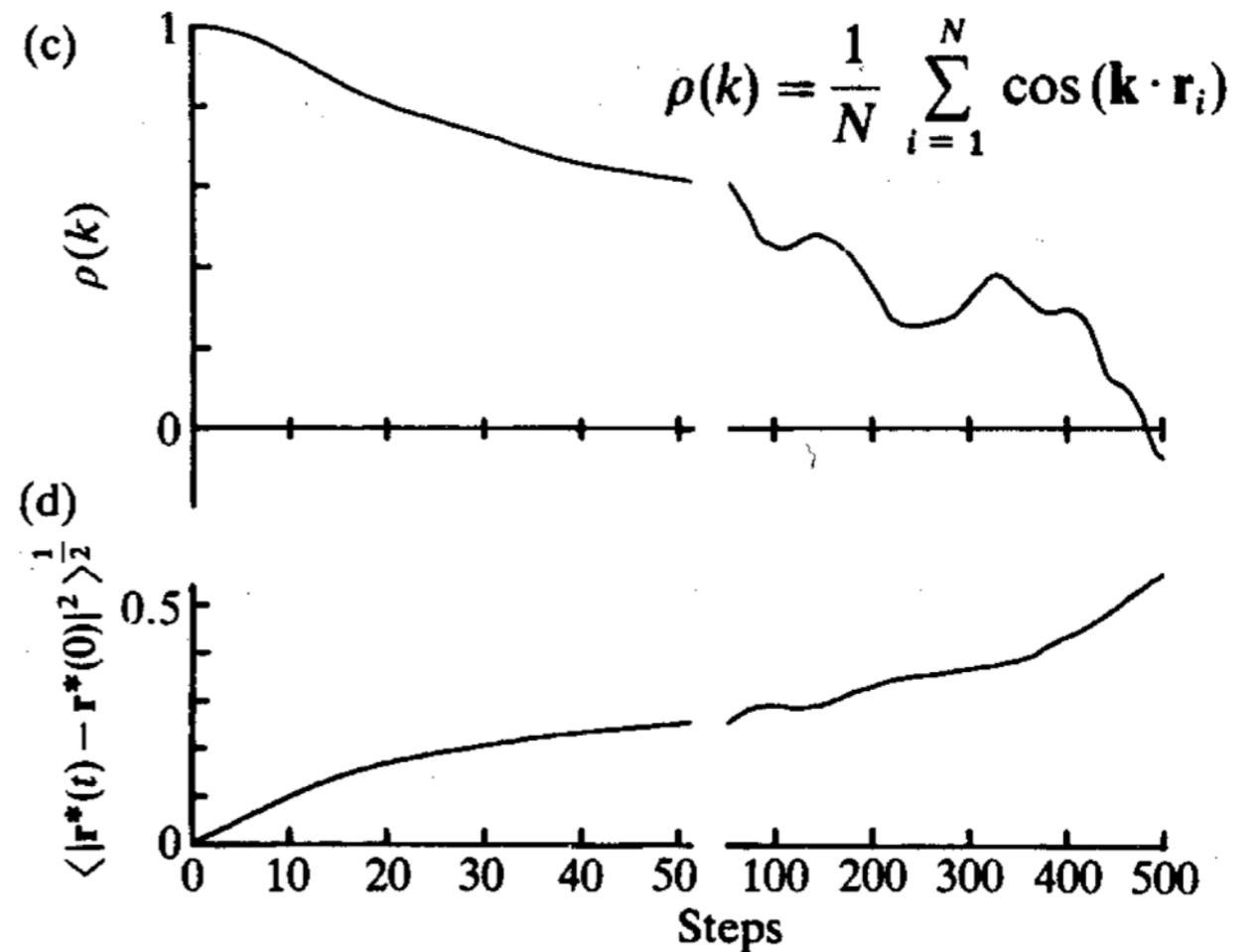
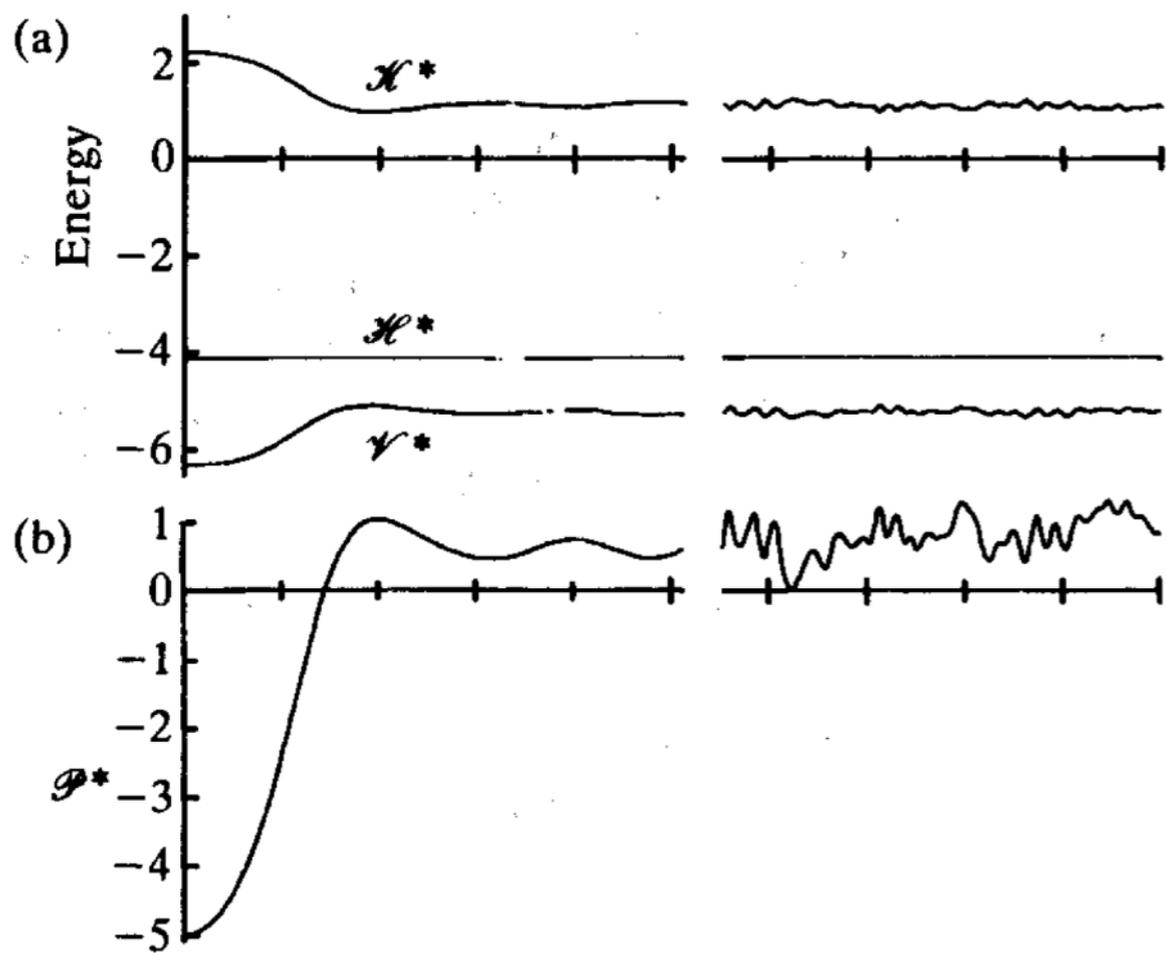
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Equilibrating the System

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- ❖ To speed equilibration from lattice, can often increase \mathbf{T} to “melt” the lattice, then later decrease to desired \mathbf{T}

More on Equilibration

- ❖ Scale interactions to speed equilibration in some cases
- ❖ Starting closer to equilibrium state results in shorter equilibration times
- ❖ Difficult to say *a priori* how long is needed for equilibration
- ❖ More time typically needed when equilibrating from a lattice or near a phase transition
- ❖ **Golden Rule:** examine carefully parameters/observables as the simulation proceeds
- ❖ Check as many exact relations as possible, e.g. partitioning of kinetic energy in molecular systems
- ❖ Once equilibrated, can proceed to **production** phase: accumulating meaningful statistics!

Simulations at Constant Temperature

- ❖ Under most experimental situations, we operate at a constant temperature T
- ❖ How can we do this in a MD simulation?

Simulations at Constant Temperature

- ❖ Under most experimental situations, we operate at a constant temperature T
- ❖ How can we do this in a MD simulation?
- ❖ Imagine the system is **weakly coupled** to a **heat bath** (thermal reservoir) at specified temperature T
- ❖ We will discuss 3 approaches:
 - Constraint Methods
 - Extended System Methods
 - Stochastic Methods

Constraint Method: Velocity-Rescaling

- ❖ Rescaling velocities at every step would yield correct **T**, but not preserve fluctuations in KE: *not a canonical ensemble!*

Constraint Method: Velocity-Rescaling

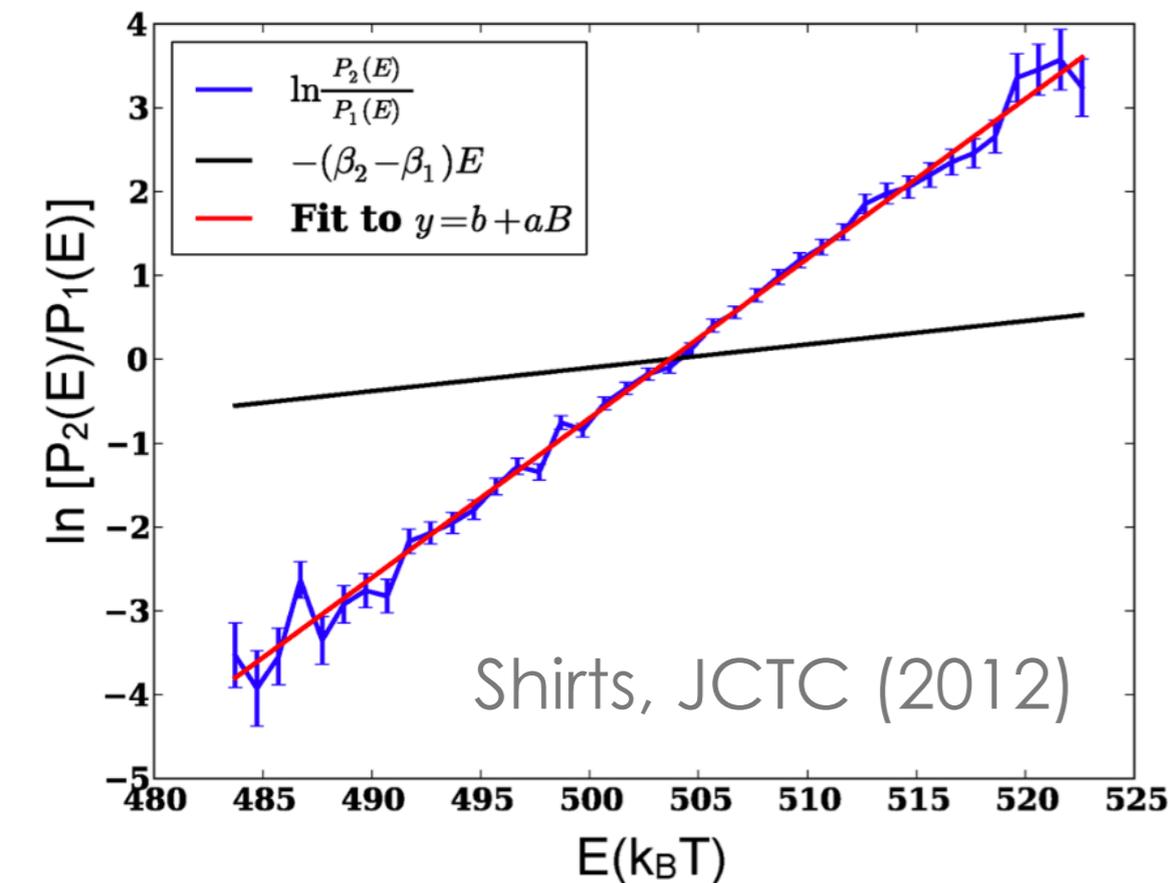
- ❖ Rescaling velocities at every step would yield correct \mathbf{T} , but not preserve fluctuations in KE: *not a canonical ensemble!*
- ❖ **Berendsen thermostat** (1984): scale over some time scale

$$\mathbf{v}_{\text{new}} = \lambda \mathbf{v} \quad \lambda^2 = 1 + \frac{\Delta t}{\tau} \left(\frac{T}{T_{\text{inst}}} - 1 \right)$$

Constraint Method: Velocity-Rescaling

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- ❖ **Berendsen does not preserve canonical distribution** (incorrect fluctuations in KE)
- ❖ Good for fast equilibration...

Canonical Velocity-Rescaling

Andersen, J. Chem. Phys. 1980; Bussi *et al.*, J. Chem. Phys. 2007

- ❖ **Andersen thermostat** & canonical velocity-rescaling yield a true canonical ensemble
- ❖ Strength of coupling to heat bath specified by collision frequency ν .
- ❖ For each particle at each step, if random number between 0 & 1 $< \nu\Delta t$, then particle velocities are reset from a Gaussian

$$P(v_x) = \left(\frac{m}{2\pi k_B T} \right)^{1/2} \exp \left(-\frac{mv_x^2}{2k_B T} \right)$$

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Andersen, J. Chem. Phys. 1980; Bussi *et al.*, J. Chem. Phys. 2007

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- ❖ **Generates canonical ensemble over long times**
- ❖ **Significantly perturbs dynamics**
- ❖ Bussi *et al.* canonical velocity-rescaling minimally affects dynamic properties

Nose-Hoover Thermostat

- ❖ **Nose (1984)**: rigorously formulated using Lagrangian mechanics (we simplify here)
- ❖ Introduce degrees of freedom for bath
 - s - “position” of bath
 - p_s - conjugate “momentum” of bath
 - Q - effective “mass” associated with s

Nose-Hoover Thermostat

❖ **Nose (1984)**: rigorously formulated using Lagrangian mechanics (we simplify here)

❖ Introduce degrees of freedom for bath

- s - “position” of bath
- p_s - conjugate “momentum” of bath
- Q - effective “mass” associated with s

$$\mathcal{H} = \sum_{i=1}^N \frac{m_i s^2 \mathbf{v}_i^2}{2} + U(\mathbf{r}^N) + \frac{p_s^2}{2Q} + k_B T (3N + 1) \ln s$$

❖ Momenta scaled by s ; coupling system to bath

❖ **Consider microcanonical simulation in extended system**

$$\Omega = \text{const} \times Q(N, V, T)$$

Nose-Hoover Thermostat

❖ **Hoover (1986)**: scaling of momenta by s is not convenient to implement; Hoover reformulated approach

❖ Introduce friction coefficient that “replaces p_s ”

$$\mathcal{H} = \sum_{i=1}^N \frac{\mathbf{p}_i^2}{2m_i} + U(\mathbf{r}^N) + \frac{\xi^2 Q}{2} + 3Nk_B T \ln s$$

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- ❖ Velocity update modified by term (friction x velocity)

- ❖ Can develop analogues of Velocity Verlet

- ❖ See Martyna, Tuckerman, Tobias, Klein, Mol. Phys. (1996) for more...

Nose-Hoover Thermostat

- ❖ **Nose-Hoover Thermostat yields the canonical ensemble**

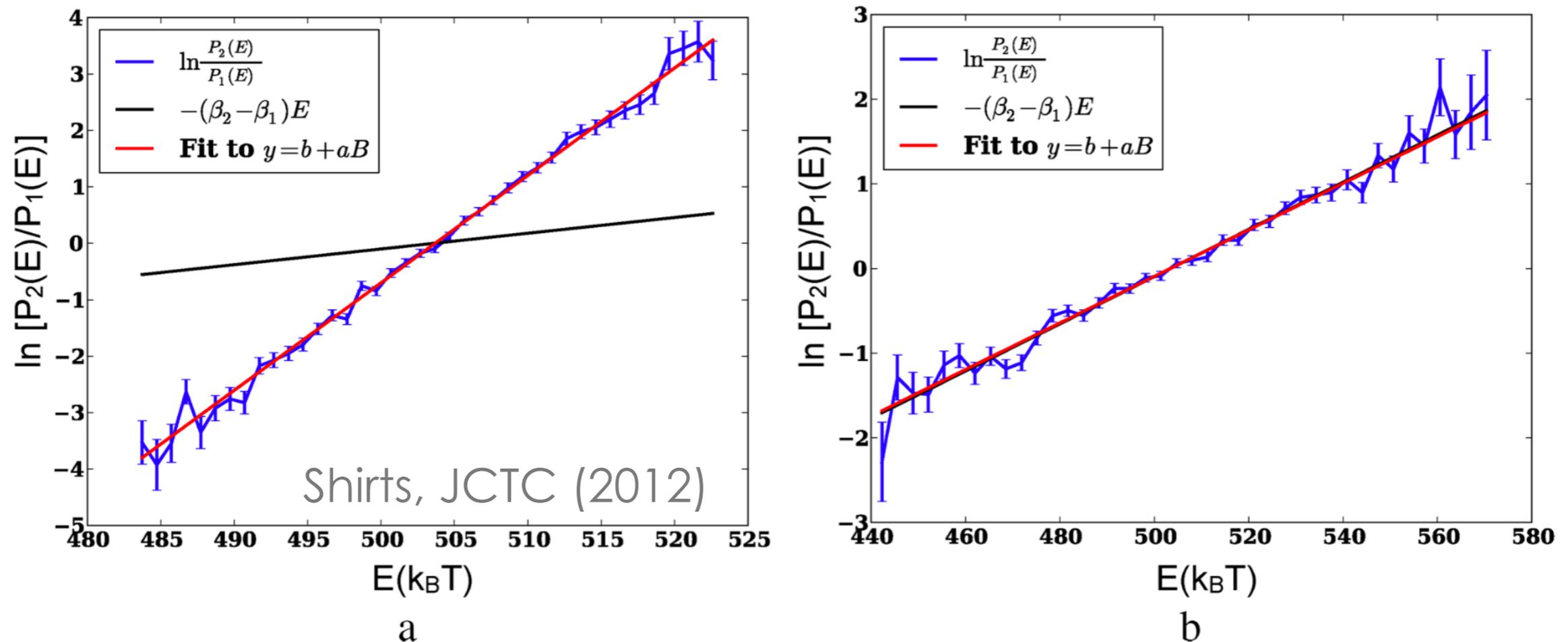


Figure 5. Differences in validation of Berendsen and Nosé–Hoover thermostats. (a) Berendsen temperature control produces simulations deviating greatly from the true distribution; in this case, the slope $\beta_2 - \beta_1$ of the kinetic energy log ratio is 7 times higher than it should be, 68 standard deviations away from the true value. (b) The Nosé–Hoover thermostat, like most others examined here, gives a slope statistically indistinguishable from the proper slope for the kinetic energy portion of the canonical ensemble.

- ❖ **Nose-Hoover chains** have been developed to increase robustness and reduce issues with NH approach...

Langevin (Stochastic) Thermostat

- ❖ Influence of the bath is captured by a **friction coefficient** and a **stochastic noise** term to mimic collisions with bath

- ❖ Modified equation of motion:

$$m_i \mathbf{a}_i = - \frac{\partial U(\mathbf{r}^N)}{\partial \mathbf{r}_i} - \gamma_i m_i \mathbf{v} + (2k_B T \gamma_i m_i)^{1/2} \eta(t)$$

- ❖ Stochastic term is Gaussian random variable: $\langle \eta(0) \eta(t) \rangle = \delta(t)$

- ❖ Rigorously conserves canonical ensemble

- ❖ Introduces stochastic component to trajectories...

Simulations at Constant Pressure

- ❖ Can readily extend thermostating approaches
- ❖ Consider coupling to “pressure bath” instead of “temperature bath”

Simulations at Constant Pressure

- ❖ Can readily extend thermostating approaches
- ❖ Consider coupling to “pressure bath” instead of “temperature bath”
- ❖ Barostats:
 - Volume rescaling (akin to velocity rescaling)
 - Berendsen barostat (doesn't produce NPT ensemble)
 - Extended ensemble: **Andersen barostat (JCP 1980)**
- ❖ Need to capture correct volume fluctuations to produce isothermal-isobaric ensemble
- ❖ Make predictions regarding density, phase transitions,...

